

When the chorus sings and laboratory plasmas join in (*)

Fulvio Zonca^{1,2}, Xin Tao³ and Liu Chen^{2,1}

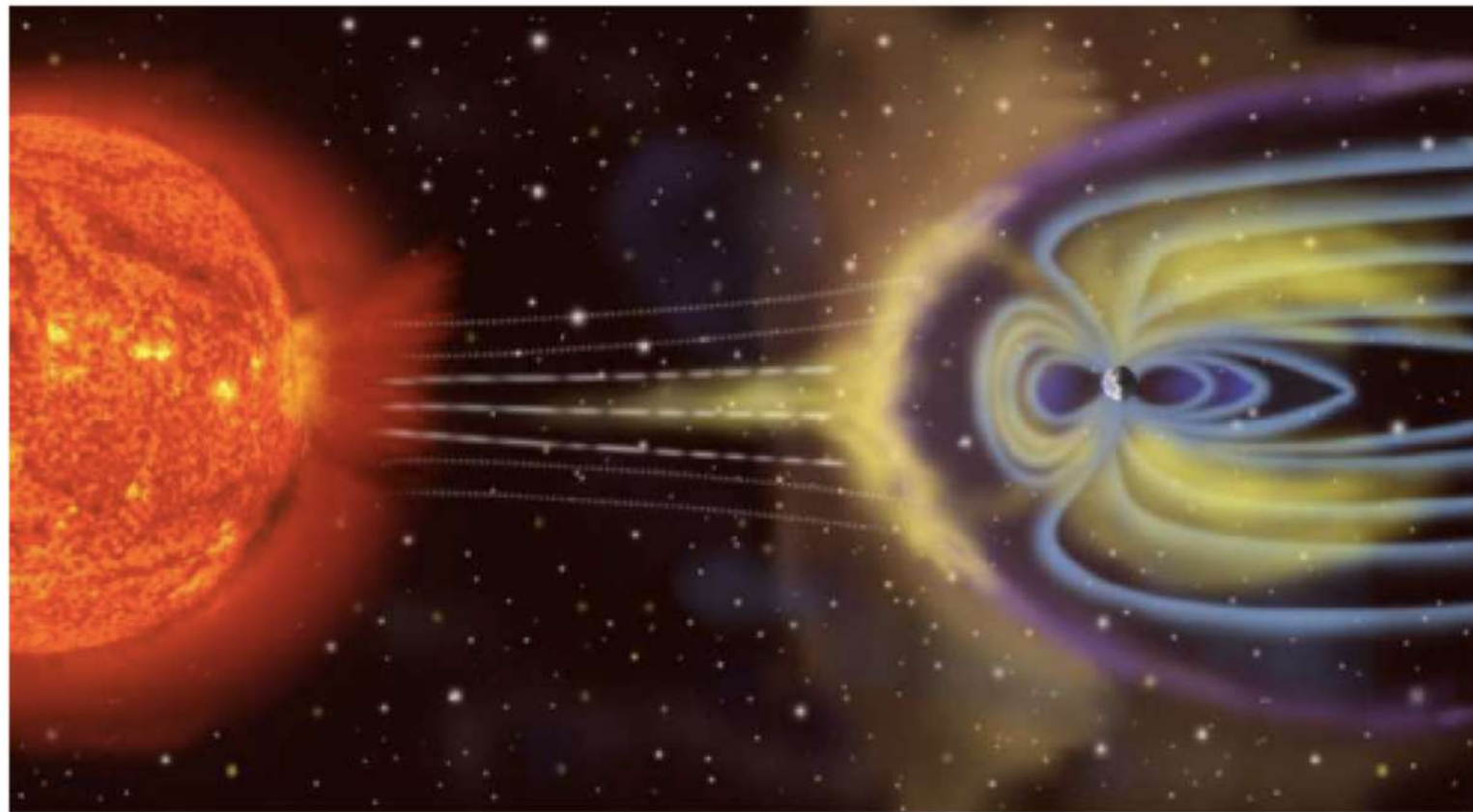
¹Center for Nonlinear Plasma Science and ENEA C.R. Frascati, 00044 Frascati, Italy

²IFTS and Dept. Physics, Zhejiang University, Hangzhou, 310027 P.R. China

³Department of Geophysics and Planetary Sciences, USTC, Hefei, P.R. China

(*) Contributions from: S. Briguglio, M.V. Falessi, G. Fogaccia, V. Fusco,
Z. Qiu, G. Vlad, X. Wang

What is chorus



Source: <http://sec.gsfc.nasa.gov/popscise.jpg>

- The Earth's magnetospheric plasma, under certain conditions, can amplify specific types of electromagnetic waves.
- In particular, the whistler waves, which are electromagnetic fluctuations propagating at less than the non-relativistic electron cyclotron frequency, can occur in $\sim 0.1s$ bursts with either rising or falling tones.

Chorus chirping

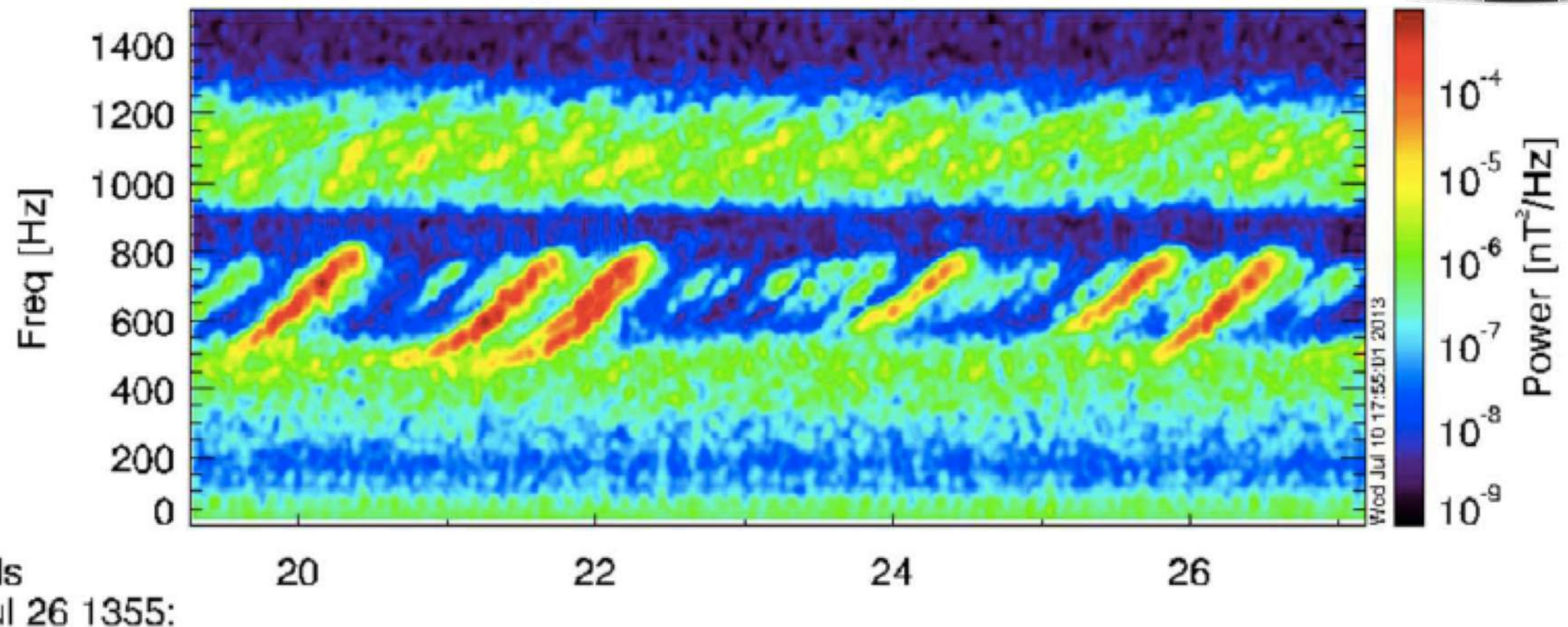
- These bursts are known as chorus because of their characteristic chirping.



[Source: <https://www.ucalgary.ca/above/science/chorus>]



Magnetic field from Themis-A



- Chorus excitation and nonlinear dynamics is one of the long-studied physics problems of Earth's magnetosphere due to its implications for particle acceleration and distribution in the radiation belts.

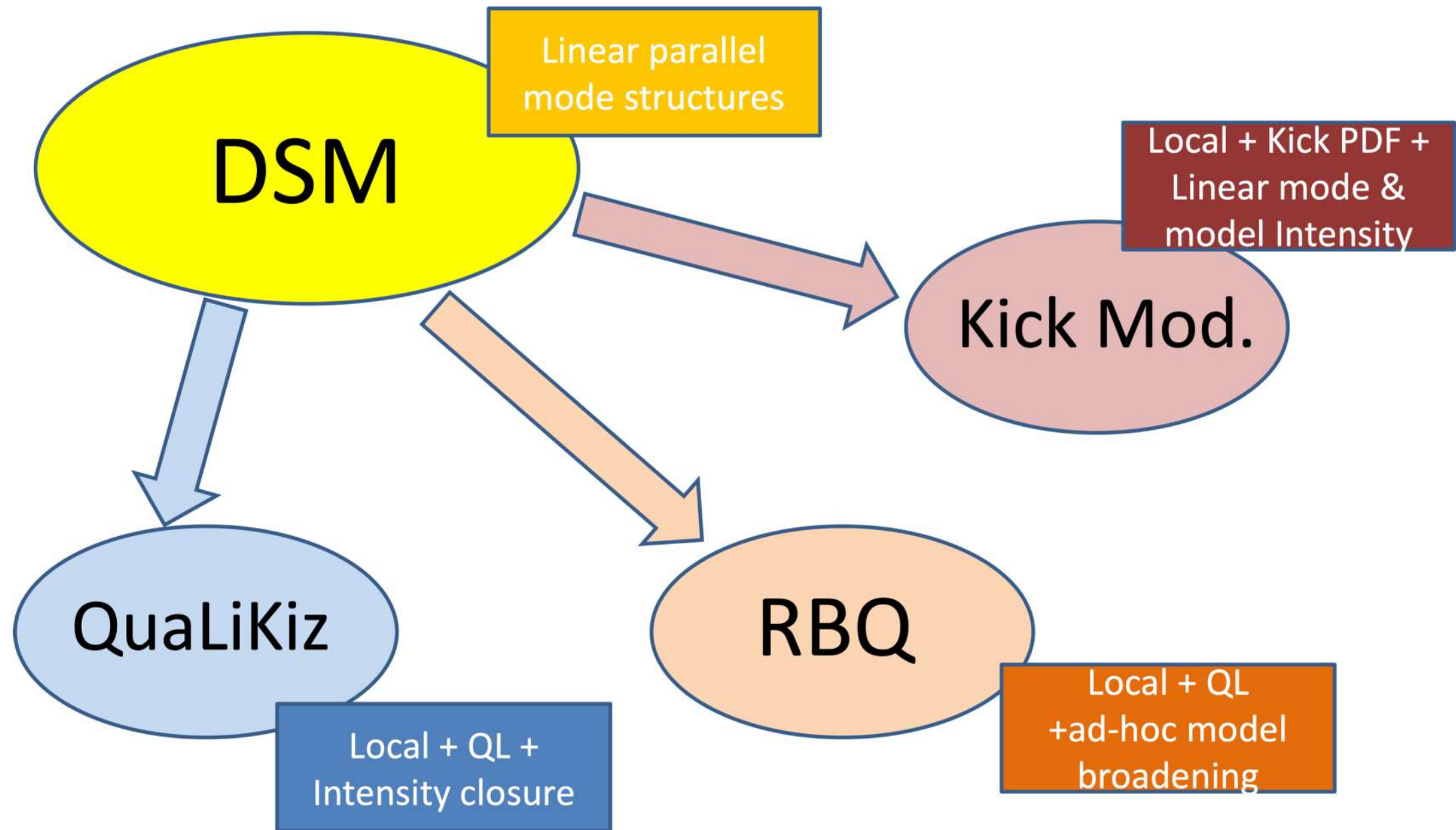
Chorus and fusion plasmas - I

- This talk: focus on (nearly) parallel propagating whistler giving rise to rising tone chorus generation.
 - Self-consistent nonlinear dynamics of non-uniform, non-autonomous system with one degree of freedom
 - Fundamental plasma physics problem of general interest and practical applications
 - Connection with reduced models with application to magnetized fusion plasmas [Chen&Zonca RMP 2016]

- Nonlinear wave particle interaction time during chorus wave generation is typically of the same order as the characteristic nonlinear time scale.
 - Nonlinear mode evolution (τ_{NL}) occurs self-consistently with particle transport (τ_{tr})
 - Novelty of the present work [Tao, Chen & Zonca 2017, 2019] with respect to previous analyses assuming $\tau_{NL} \gg \tau_{tr}$

Chorus and fusion plasmas - II

Direct connection with Dyson Schrödinger Model (DSM) :
(ZCFQ seminar Nov 20/2020)



- **Numerical simulation results summarized:**
 - A “Trap-Release-Amplify” Model of Chorus Waves, published in Journal of Geophysical Research: Space Physics ([PDF](#))
- **Theoretical framework fully analyzed:**
 - A theoretical framework of chorus wave excitation (arXiv preprint arXiv:2107.03151; submitted to JGR; [PDF](#))
 - Nonlinear dynamics and phase space transport by chorus emission (arXiv preprint arXiv:2106.15026; tutorial submitted to RMPP; [PDF](#))

Wave packet evolution - I

- Evolution equation for the fluctuation spectrum using wave-packets
 - $\delta \mathbf{E}_{\perp}(z, t) = e^{iS_k(z, t)} \delta \bar{\mathbf{E}}_{\perp k}(z, t); \omega_k = -\partial_t S_k, k = \partial_z S_k$
 - $\delta \bar{\mathbf{E}}_{\perp k}(z, t) = |\delta \bar{\mathbf{E}}_{\perp k}(z, t)| e^{i\varphi_k(z, t)}; I_k(z, t) \equiv (\partial D_w / \partial k) |\delta \mathbf{E}_{\perp 0k}(z, t)|^2$

- Wave packets satisfy the Whistler wave dispersion relation

$$D_w = 1 + \frac{\omega_p^2}{\omega(\Omega - \omega)} - \frac{k^2 c^2}{\omega^2} = 0 \quad \Rightarrow \quad k = K(z, \omega)$$

$$\left(1 + \frac{\omega_p^2}{\omega(\Omega - \omega)} \right) \delta \bar{\mathbf{E}}_{\perp k} - \frac{k^2 c^2}{\omega^2} \delta \bar{\mathbf{E}}_{\perp k} = -\frac{4\pi i}{\omega} \delta \bar{\mathbf{J}}_{hk}$$

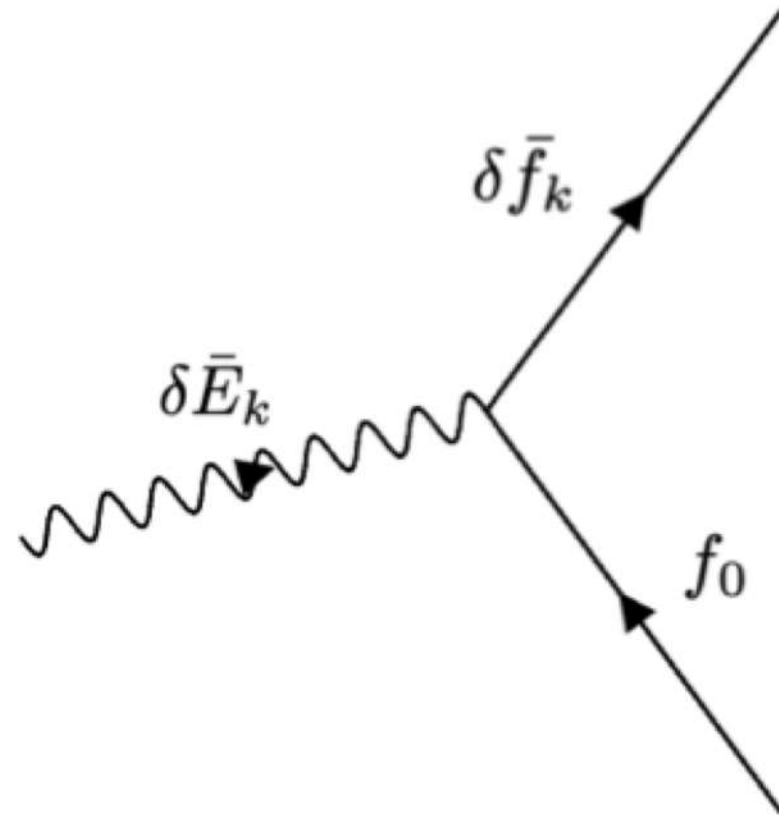
- Wave packet propagation is entirely constructed from

$$W(z, t, \omega) + i\Gamma(z, t, \omega) \equiv - \frac{4\pi i}{\omega_k \partial D_w / \partial \omega_k} \frac{\delta \mathbf{J}_h \cdot \delta \mathbf{E}_{\perp}^*}{|\delta \mathbf{E}_{\perp}(z, t)|^2} \Big|_{k=K(z, \omega)}$$

$\delta \mathbf{J}_h \equiv$ supra-thermal electron current

Linear electron response

- Compute electron current from particle response:



$$\delta \bar{f}_k = \frac{e}{m} \mathcal{P}_k v_{\perp} \delta \bar{E}_k \bar{\partial}_{\mathcal{E}k} f_0 ,$$

Electron response

$$\begin{aligned} \bar{\partial}_{\mathcal{E}k} &\equiv \left[\frac{k}{\omega_k} \frac{\partial}{\partial v_{\parallel}} + \left(1 - \frac{k v_{\parallel}}{\omega_k} \right) \frac{1}{v_{\perp}} \frac{\partial}{\partial v_{\perp}} \right] \\ &= \left[\frac{\partial}{\partial \mathcal{E}} + \left(1 - \frac{k v_{\parallel}}{\omega_k} \right) \frac{1}{B} \frac{\partial}{\partial \mu} \right] ; \end{aligned}$$

$$\delta \bar{\mathbf{J}}_{hk\perp} = -e \langle \mathbf{v}_{\perp} e^{i\alpha} \delta \bar{f}_k \rangle$$

$$\mathcal{P}_k = [\Omega + k v_{\parallel} - \omega_k]^{-1}$$

Linear propagator

$$\frac{4\pi i}{\omega} \frac{\delta \bar{\mathbf{J}}_{hk} \cdot \delta \bar{\mathbf{E}}_{\perp k}^*}{|\delta \bar{\mathbf{E}}_{\perp k}(z, t)|^2} \Big|_{k=K(z, \omega)} = \left\langle \frac{4\pi e}{\omega} \frac{v_{\perp}}{2} \frac{\delta \bar{E}_k^* \delta \bar{f}_k}{|\delta \bar{E}_k|^2} \right\rangle$$

$$= \frac{\omega_p^2}{n\omega} \left\langle \frac{v_{\perp}^2/2}{\Omega + k v_{\parallel} - \omega} \left[\frac{k}{\omega} \frac{\partial f_0}{\partial v_{\parallel}} + \left(1 - \frac{k v_{\parallel}}{\omega} \right) \frac{1}{v_{\perp}} \frac{\partial f_0}{\partial v_{\perp}} \right] \right\rangle$$

Wave packet evolution - II

- Intensity evolution equation

$$\left(\frac{\partial}{\partial t} + v_g \frac{\partial}{\partial z} \right) I_k(z, t) = 2\Gamma(z, t, \omega) I_k(z, t)$$

- Phase evolution equation

$$\left(\frac{\partial}{\partial t} + v_g \frac{\partial}{\partial z} \right) \varphi_k(z, t) = -W(z, t, \omega)$$

- As noted already, all relevant physics is included in the wave-particle interaction term:

$$W(z, t, \omega) + i\Gamma(z, t, \omega) \equiv - \frac{4\pi i}{\omega_k \partial D_w / \partial \omega_k} \frac{\delta \mathbf{J}_h \cdot \delta \mathbf{E}_\perp^*}{|\delta \mathbf{E}_\perp(z, t)|^2} \Big|_{k=K(z, \omega)}$$

- It must be computed self-consistently, in the presence of evolving finite-amplitude fluctuation spectrum:

$\Rightarrow \tau_{NL} \sim \tau_{tr}$ (unlike existing theories that assume $\tau_{NL} \gg \tau_{tr}$)

$\Rightarrow$ Renormalization problem for the long wavelength (spatial averaged) particle response

Self consistent NL description

- **Basic underlying idea:**

- We focus on the nonlinear particle response without fast temporal or spatial dependences, which correspond to the self-interaction of the wavenumber of interest with itself. (*)
- This nonlinear response obeys an evolution equation, which may look like a “quasilinear” description at first glance but is, indeed, much more general as will be shown and discussed in the following.
- The present approach has a close connection of with the Dyson equation and the so-called “principal series” of secular terms.

(*) Acknowledgment: D. F. Escande

Holes and clumps

- Formation of structures in phase space:**

- For rising tone chorus, phase locked particles gain energy and momentum but loose parallel energy

$$\dot{\mu}B\omega = \Omega\dot{\mathcal{E}}, \quad \Rightarrow \quad \omega d\mu = (\Omega/B)d\mathcal{E}.$$

- Structures are formed near resonance:

Holes: lower density w.r.t. unperturbed surrounding

Clumps: higher density w.r.t. unperturbed surrounding

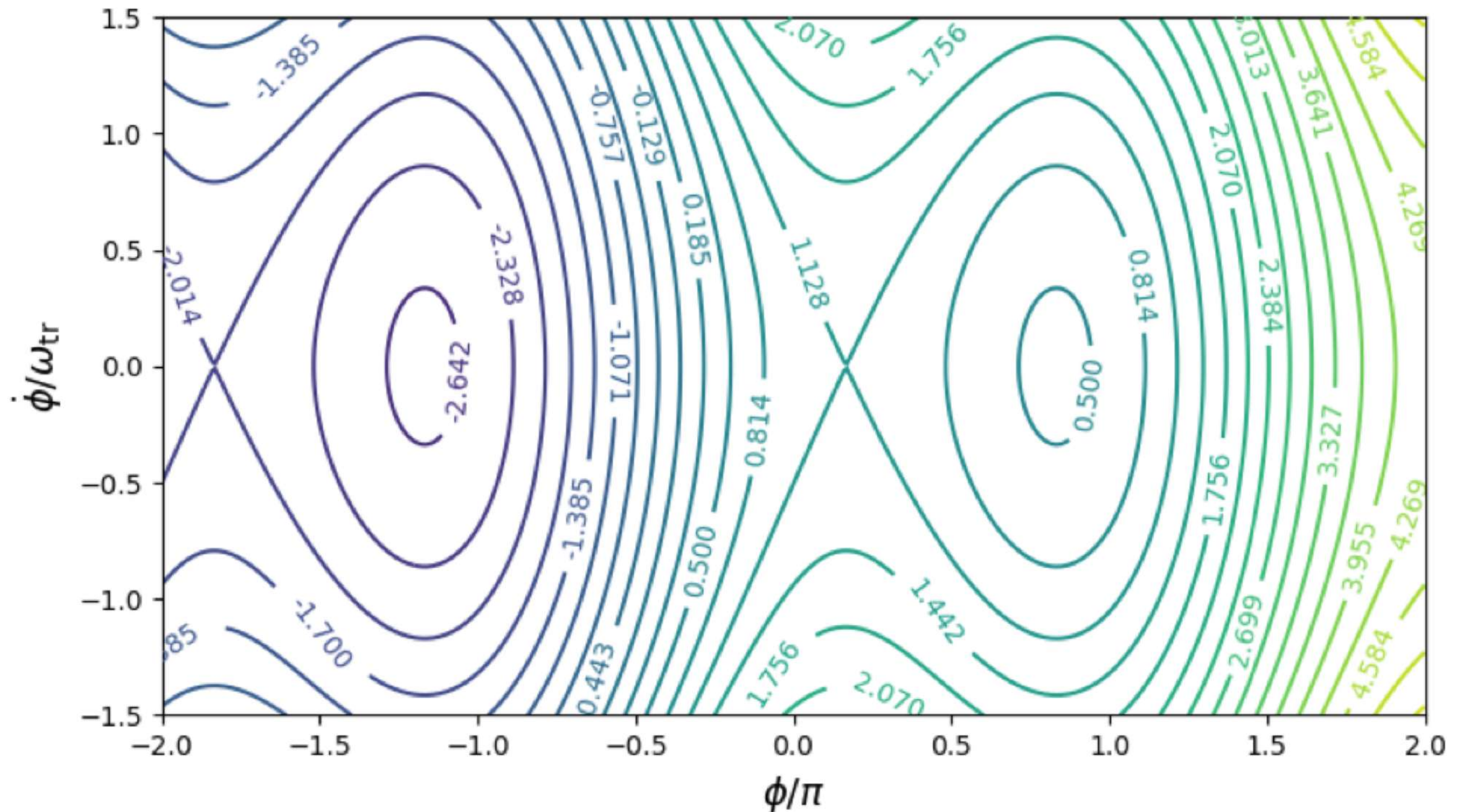
$$f_0 = \frac{n_e}{(2\pi)^{3/2}w_{\parallel e}w_{\perp e}^2} \exp\left(-\mathcal{E}/w_{\parallel e}^2 + A\mu B_e/w_{\perp e}^2\right),$$

$$A = w_{\perp e}^2/w_{\parallel e}^2 - 1$$

$$\Delta f_0 = f_0(\mathcal{E}, \mu) - f_0(\mathcal{E} + \Delta\mathcal{E}, \mu + \Delta\mu),$$

$$\propto \frac{\Delta\mathcal{E}}{w_{\parallel e}^2} - AB_e \frac{\Delta\mu}{w_{\perp e}^2} = -AB_e \frac{\Delta\mu}{w_{\parallel e}^2} \left(\frac{A}{A+1} - \frac{\omega}{\Omega_e} \right).$$

Holes and clumps



$$\ddot{\phi} = \omega_{tr}^2 (\sin \phi - R)$$

$$(1/2)\dot{\phi}^2 + \omega_{tr}^2 (\cos \phi + R\phi) = \text{const} \quad R = \frac{1}{\omega_{tr}^2} \left[\left(1 - \frac{v_r}{v_g}\right)^2 \frac{\partial \omega}{\partial t} + \left(\frac{kv_{\perp}^2}{2\Omega} - \frac{3}{2}v_r\right) \frac{\partial \Omega}{\partial z} \right]$$

- **Asymptotic expansion:**

- Formally solving NL Vlasov using small amplitude expansion

$$f = f^{(0)} + f^{(1)} + f^{(2)} + \dots$$

$$\left(\frac{\partial}{\partial t} + v_{\parallel} \frac{\partial}{\partial z} + \Omega \frac{\partial}{\partial \alpha} \right) f - \frac{e}{m} \left(\delta \mathbf{E}_{\perp} + \frac{\mathbf{v} \times \delta \mathbf{B}_{\perp}}{c} \right) \cdot \frac{\partial}{\partial \mathbf{v}} f = 0$$

$$f(\mathbf{v}, z, t) = f_0(\mathbf{v}, z) + \frac{1}{2} \sum_k \left(e^{iS_k(z,t) + i\alpha} \delta \bar{f}_k(\mathbf{v}, z, t) + c.c. \right)$$

$$\begin{aligned} f^{(2)} = & \frac{1}{2} \sum_{k,k'} \left(e^{iS_k(z,t) + iS_{k'}(z,t) + 2i\alpha} \delta \bar{f}_{k+k'}(\mathbf{v}, z, t) + c.c. \right) \\ & + \frac{1}{2} \sum_{k,k'} \left(e^{iS_k(z,t) - iS_{k'}(z,t)} \delta \bar{f}_{k-k'}(\mathbf{v}, z, t) + c.c. \right) . \end{aligned}$$

- Asymptotic expansion:**

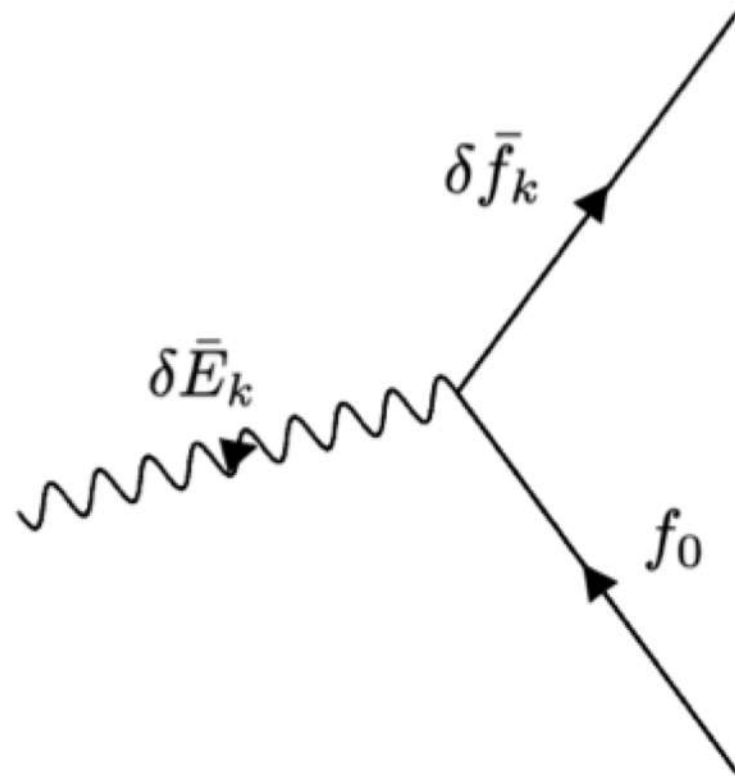
- Formally solving NL Vlasov using small amplitude expansion

$$\left[2\Omega + (k + k')v_{\parallel} - (\omega_k + \omega_{k'}) - i(\partial_t + v_{\parallel}\partial_z) \right] \delta \bar{f}_{k+k'} = \frac{e}{m} \frac{v_{\perp}}{4} \left\{ \delta \bar{E}_k \left[\bar{\partial}_{\mathcal{E}k} - \frac{1}{v_{\perp}^2} \left(1 - \frac{kv_{\parallel}}{\omega_k} \right) \right] \delta \bar{f}_{k'} + \delta \bar{E}_{k'} \left[\bar{\partial}_{\mathcal{E}k'} - \frac{1}{v_{\perp}^2} \left(1 - \frac{k'v_{\parallel}}{\omega_{k'}} \right) \right] \delta \bar{f}_k \right\},$$

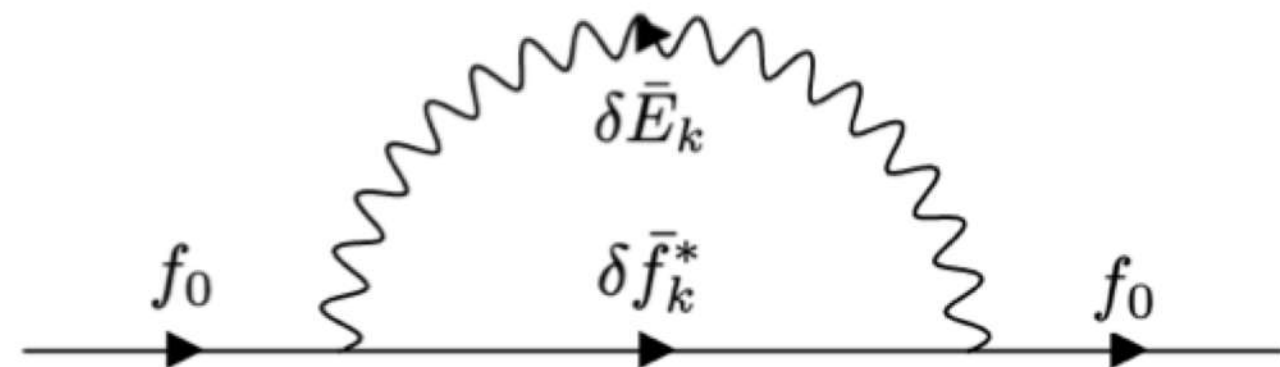
$$\begin{aligned} \bar{\partial}_{\mathcal{E}k} &\equiv \left[\frac{k}{\omega_k} \frac{\partial}{\partial v_{\parallel}} + \left(1 - \frac{kv_{\parallel}}{\omega_k} \right) \frac{1}{v_{\perp}} \frac{\partial}{\partial v_{\perp}} \right] \\ &= \left[\frac{\partial}{\partial \mathcal{E}} + \left(1 - \frac{kv_{\parallel}}{\omega_k} \right) \frac{1}{B} \frac{\partial}{\partial \mu} \right]; \end{aligned}$$

$$\begin{aligned} \left[(k - k')v_{\parallel} - (\omega_k - \omega_{k'}) - i(\partial_t + v_{\parallel}\partial_z) \right] \delta \bar{f}_{k-k'} &= \frac{e}{m} \frac{v_{\perp}}{4} \left\{ \delta \bar{E}_k \left[\bar{\partial}_{\mathcal{E}k} + \frac{1}{v_{\perp}^2} \left(1 - \frac{kv_{\parallel}}{\omega_k} \right) \right] \delta \bar{f}_{k'}^* - \delta \bar{E}_{k'}^* \left[\bar{\partial}_{\mathcal{E}k'} + \frac{1}{v_{\perp}^2} \left(1 - \frac{k'v_{\parallel}}{\omega_{k'}} \right) \right] \delta \bar{f}_k \right\}. \end{aligned} \quad (52)$$

- **Renormalization via same-k interactions:**
 - Consider the NL contribution without fast spatiotemporal scales
 - Effect of self-interactions of the wave number of interest with itself



(a)



(b)

- Renormalization via same-k interactions:

- Secular growth on short time scale

$$f_0 \rightarrow f_0 + \sum_k \text{Re} \delta \bar{f}_0 ,$$

$$(\partial_t + v_{\parallel} \partial_z) \delta \bar{f}_0 = \frac{e}{m} \frac{v_{\perp}}{4} \sum_k i \left\{ \delta \bar{E}_k \left[\bar{\partial}_{\varepsilon k} + \frac{1}{v_{\perp}^2} \right. \right. \\ \times \left. \left. \left(1 - \frac{k v_{\parallel}}{\omega_k} \right) \right] \delta \bar{f}_k^* - \delta \bar{E}_k^* \left[\bar{\partial}_{\varepsilon k} + \frac{1}{v_{\perp}^2} \left(1 - \frac{k v_{\parallel}}{\omega_k} \right) \right] \delta \bar{f}_k \right\}$$

$$\delta \bar{f}_k = \frac{e}{m} \mathcal{P}_k v_{\perp} \delta \bar{E}_k \bar{\partial}_{\varepsilon k} f_0 ,$$

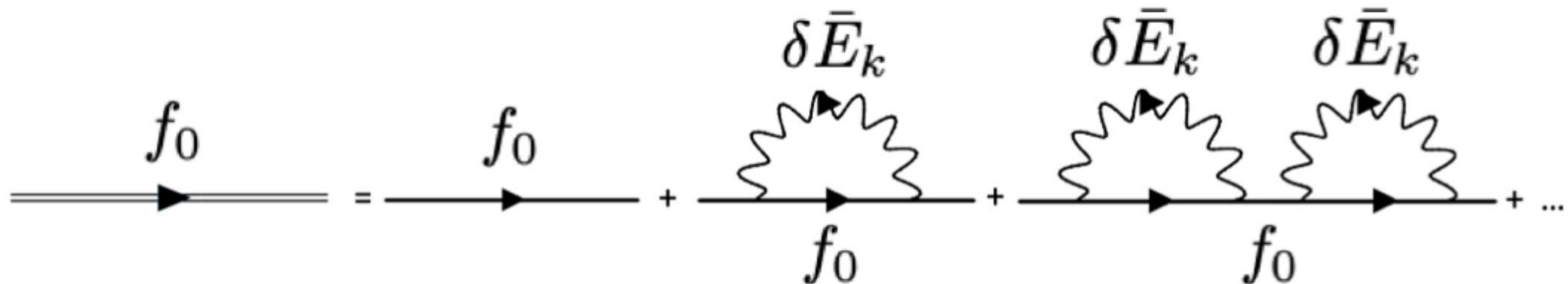
$$\mathcal{P}_k = [\Omega + k v_{\parallel} - \omega_k]^{-1}$$



- Renormalization via same-k interactions:

- By iterative solution one obtains the dominant renormalized response → **Dyson – like equation (DSE)**

$$(\partial_t + v_{\parallel} \partial_z) f_0 = \frac{e}{m} \frac{v_{\perp}}{4} \text{Re} \sum_k i \left\{ \delta \bar{E}_k \left[\bar{\partial}_{\varepsilon k} + \frac{1}{v_{\perp}^2} \left(1 - \frac{kv_{\parallel}}{\omega_k} \right) \right] \delta \bar{f}_k^* - \delta \bar{E}_k^* \left[\bar{\partial}_{\varepsilon k} + \frac{1}{v_{\perp}^2} \left(1 - \frac{kv_{\parallel}}{\omega_k} \right) \right] \delta \bar{f}_k \right\}$$



$$\delta \bar{f}_k = \frac{e}{m} \mathcal{P}_k v_{\perp} \delta \bar{E}_k \bar{\partial}_{\varepsilon k} f_0 ,$$

$$\mathcal{P}_k \equiv \left[\Omega + kv_{\parallel} - \omega_k - i(\partial_t + v_{\parallel} \partial_z) - \Delta_k \right]^{-1}$$

resonance broadening



浙江大学聚变理论模拟中心 潘云鹤

Institute for Fusion Theory and Simulation, Zhejiang University

Resonance Broadening

- Coupling of $\delta \bar{f}_{k+k'}$ with $(k')^*$ and $\delta \bar{f}_{k-k'}$ with k' (Dupree 66):

$$\delta \bar{f}_k = \frac{e}{m} \mathcal{P}_k v_{\perp} \delta \bar{E}_k \bar{\partial}_{\mathcal{E}k} f_0 ,$$

$$\mathcal{P}_k \equiv [\Omega + kv_{\parallel} - \omega_k - i(\partial_t + v_{\parallel} \partial_z) - \Delta_k]^{-1} \quad \Delta_k \equiv \Delta_{1k} + \Delta_{2k}$$

$$\bar{\partial}_{\mathcal{E}k} \equiv \left[\frac{k}{\omega_k} \frac{\partial}{\partial v_{\parallel}} + \left(1 - \frac{kv_{\parallel}}{\omega_k} \right) \frac{1}{v_{\perp}} \frac{\partial}{\partial v_{\perp}} \right]$$

$$= \left[\frac{\partial}{\partial \mathcal{E}} + \left(1 - \frac{kv_{\parallel}}{\omega_k} \right) \frac{1}{B} \frac{\partial}{\partial \mu} \right] ;$$

$$\begin{aligned} -i\Delta_{1k} &\equiv i \frac{e^2}{m^2} \frac{v_{\perp}^2}{8} \sum_{k'} \delta \bar{E}_{k'}^* \left[\bar{\partial}_{\mathcal{E}k'} + \frac{2}{v_{\perp}^2} \left(1 - \frac{k'v_{\parallel}}{\omega_{k'}} \right) \right] \\ &\quad \times \mathcal{P}_{k+k'} \delta \bar{E}_{k'} \left[\bar{\partial}_{\mathcal{E}k'} - \frac{1}{v_{\perp}^2} \left(1 - \frac{k'v_{\parallel}}{\omega_{k'}} \right) \right] , \end{aligned}$$

$$-i\Delta_{2k} \equiv i \frac{e^2}{m^2} \frac{v_{\perp}^2}{8} \sum_{k' \neq k} \delta \bar{E}_{k'}^* \bar{\partial}_{\mathcal{E}k'} \mathcal{P}_{k-k'} \delta \bar{E}_{k'} \left[\bar{\partial}_{\mathcal{E}k'} + \frac{1}{v_{\perp}^2} \left(1 - \frac{k'v_{\parallel}}{\omega_{k'}} \right) \right]$$

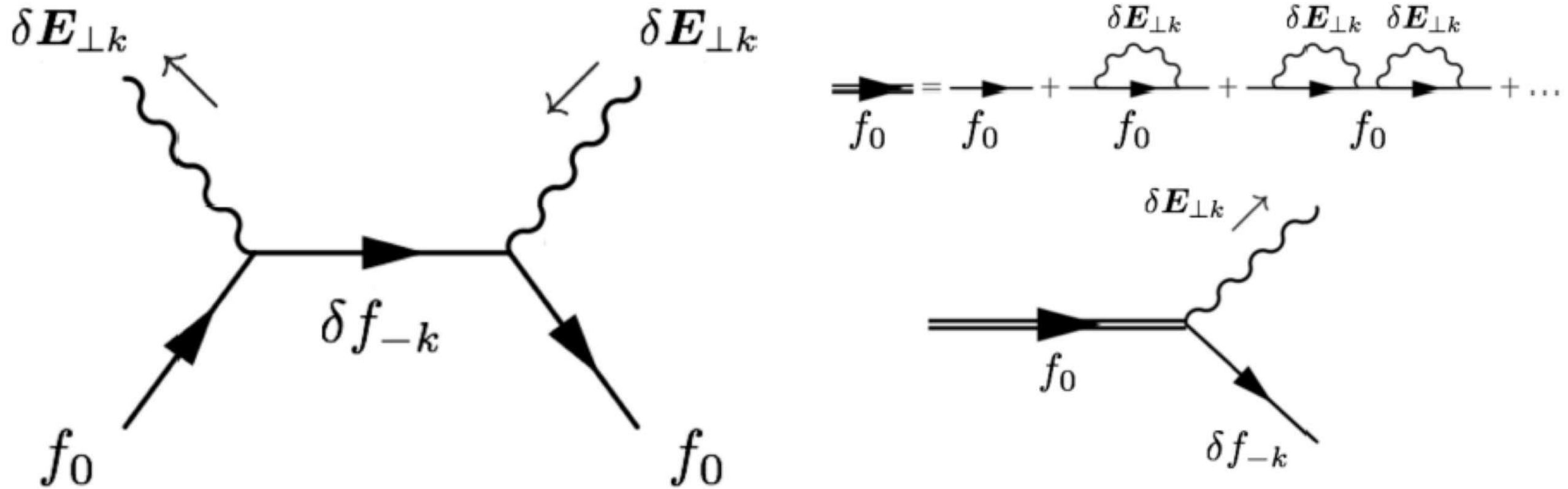
resonance broadening

Brief recap :

- Perturbation expansion of particle response in the weak field intensity limit, yields secular terms dominated by emission and absorption of the same k -fluctuation (called diagonal interactions).
- At ℓ -th order, distortion of the reference state, f_0 contains, $\propto t^\ell$ secularity, while k_0 fluctuation contains $\propto t^\ell e^{-i\omega_{k_0}t}$ behavior.
- For growing modes, $\propto t^\ell$ is replaced by $\propto (\omega_k/\gamma_k)^\ell$ behavior [Montgomery 63], [Al'tshul' & Karpman 65], [Mima 73]
 - $\Rightarrow$ rising tone chorus has a quasi-coherent fluctuation spectrum with $|\gamma_k/\omega_k| \ll 1$, with a typical nonlinear evolution time, $\tau_{NL} \sim \gamma_k^{-1}$
 - $\Rightarrow$ for quasi-coherent fluctuation spectrum, particle response is obtained as asymptotic expansion in $|\omega_k|^{-1} \tau_{NL}^{-1} \ll 1$ [Chen & Zonca RMP16]

Renormalized particle response

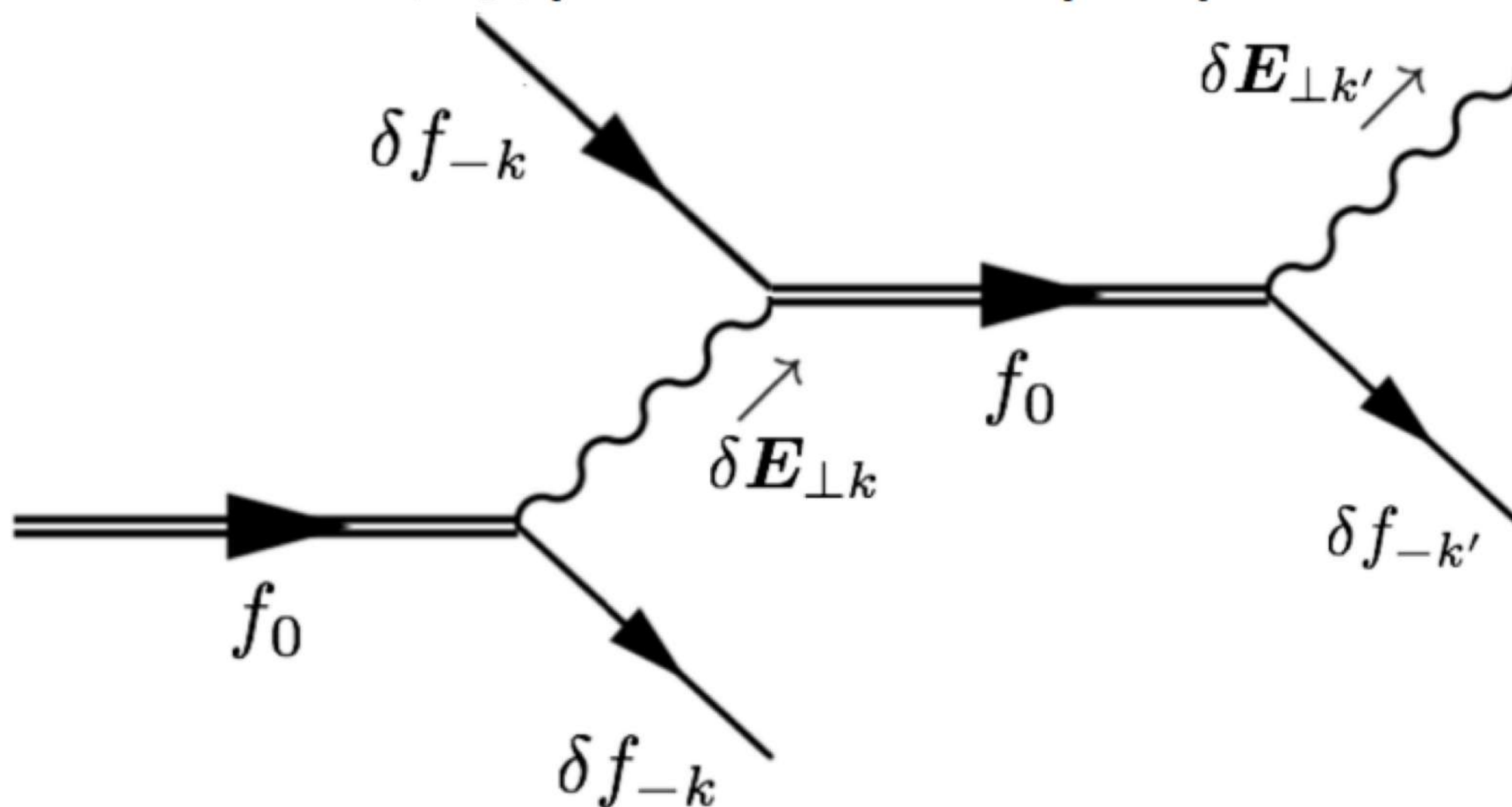
- Nonlinear dynamics takes place through the self-consistent non-linear evolution of the supra-thermal particle distribution function $f_0(t)$, which fully accounts for the supra-thermal electron current $\delta \mathbf{J}_h$ (closes self-consistent nonlinear equations). [Zonca, Tao & Chen 2017]



- The renormalized particle response is obtained via a Dyson-like equation approach.

Dyson description of chorus chirping

- Within this theoretical framework, the rising tone chorus chirping is due to the subsequent emission and reabsorption of wave-packets that maximize wave-particle power extraction [Zonca, Tao & Chen 2017]
⇒ Consistent with, e.g., [Vomvoridis et al. 1982] and [Omura et al. 2008].

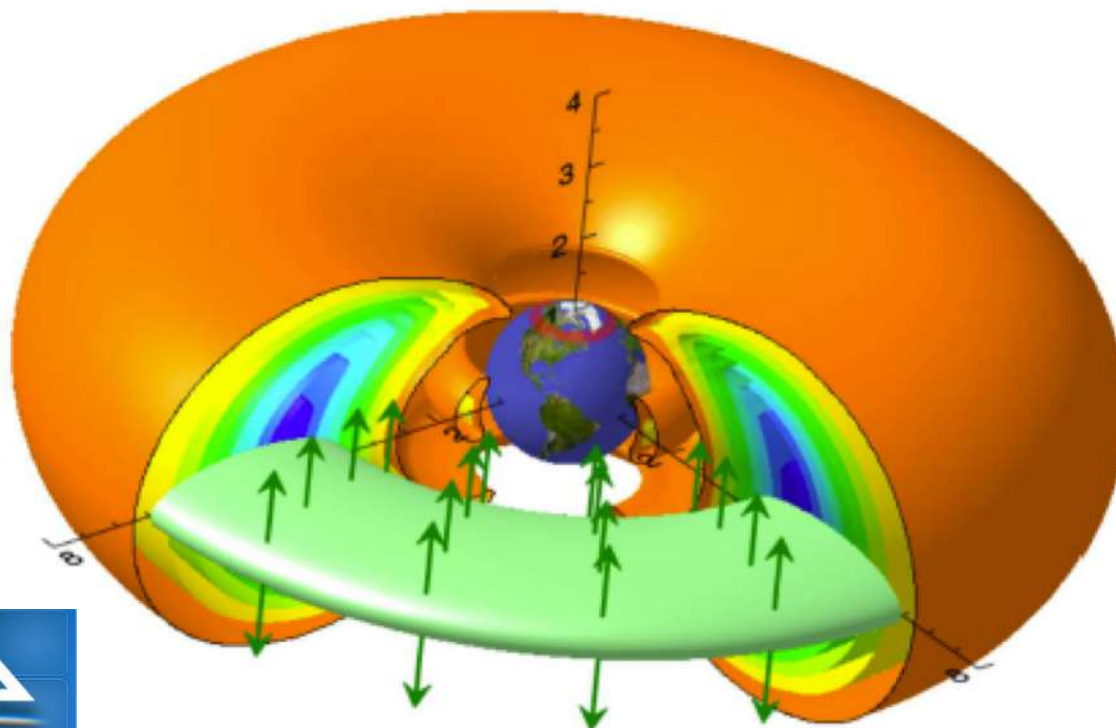


Reduced Dyson description

- Reduced Dyson description:**

- Derive evolution equation for nonlinear growth and frequency shift
- Based on the existence of a narrow (quasi-coherent) spectrum and its resonant excitation
- Use localized source to ignore magnetic field non-uniformity

$$W(z, t, \omega) + i\Gamma(z, t, \omega) = \frac{\omega(\Omega_e - \omega)^2}{\Omega_e} \frac{n_e}{n} \zeta^4 \left\langle \frac{v_{\perp}^2}{2} \mathcal{P}_k \left[\frac{kv_{\perp}^2}{2\omega} \frac{\partial \hat{f}_0}{\partial v_{\parallel}} - \frac{\Omega_e}{\omega} \hat{f}_0 \right]_{k=K(z=0, \omega)} \right\rangle$$



$$\delta \bar{f}_k = \frac{e}{m} \mathcal{P}_k v_{\perp} \delta \bar{E}_k \bar{\partial}_{\varepsilon k} f_0 ,$$

$$\mathcal{P}_k \equiv [\Omega + kv_{\parallel} - \omega_k - i(\partial_t + v_{\parallel} \partial_z) - \Delta_k]^{-1}$$

propagator

resonance broadening

Reduced Dyson description

- Reduced Dyson description:**

$$\bar{\Gamma}_{NL}(\omega) = \left[\Omega_e \frac{\partial}{\partial \omega} - \frac{2\Omega_e^2}{k^2 \langle v_{\perp}^2 \rangle} \left(1 - \frac{v_{r\omega}}{v_{g\omega}} \right) \right] \Omega_e \frac{\partial}{\partial \omega} \times \sum_{k'} \left[\frac{(\omega' - \omega)^2}{4\Omega_e^2} + \Omega_e^{-2} \partial_t^2 \right]^{-1} \times \frac{\langle \langle \omega_{trk'}^4 \rangle \rangle}{4\Omega_e^4 (1 - v_{r\omega'}/v_{g\omega'})^4} \left(\frac{\bar{\Gamma}(\omega) + \bar{\Gamma}(\omega')}{2} \right).$$

$$\Omega_e^{-1} \partial_t \bar{W}(\omega) = \left[\Omega_e \frac{\partial}{\partial \omega} - \frac{2\Omega_e^2}{k^2 \langle v_{\perp}^2 \rangle} \left(1 - \frac{v_{r\omega}}{v_{g\omega}} \right) \right] \Omega_e \frac{\partial}{\partial \omega} \times \sum_{k'} \frac{(\omega' - \omega)}{2\Omega_e} \left[\frac{(\omega' - \omega)^2}{4\Omega_e^2} + \Omega_e^{-2} \partial_t^2 \right]^{-1} \times \frac{\langle \langle \omega_{trk'}^4 \rangle \rangle}{4\Omega_e^4 (1 - v_{r\omega'}/v_{g\omega'})^4} \left(\frac{\bar{\Gamma}(\omega) + \bar{\Gamma}(\omega')}{2} \right),$$

- Narrow spectrum: recovers Vomvoridis 1982**

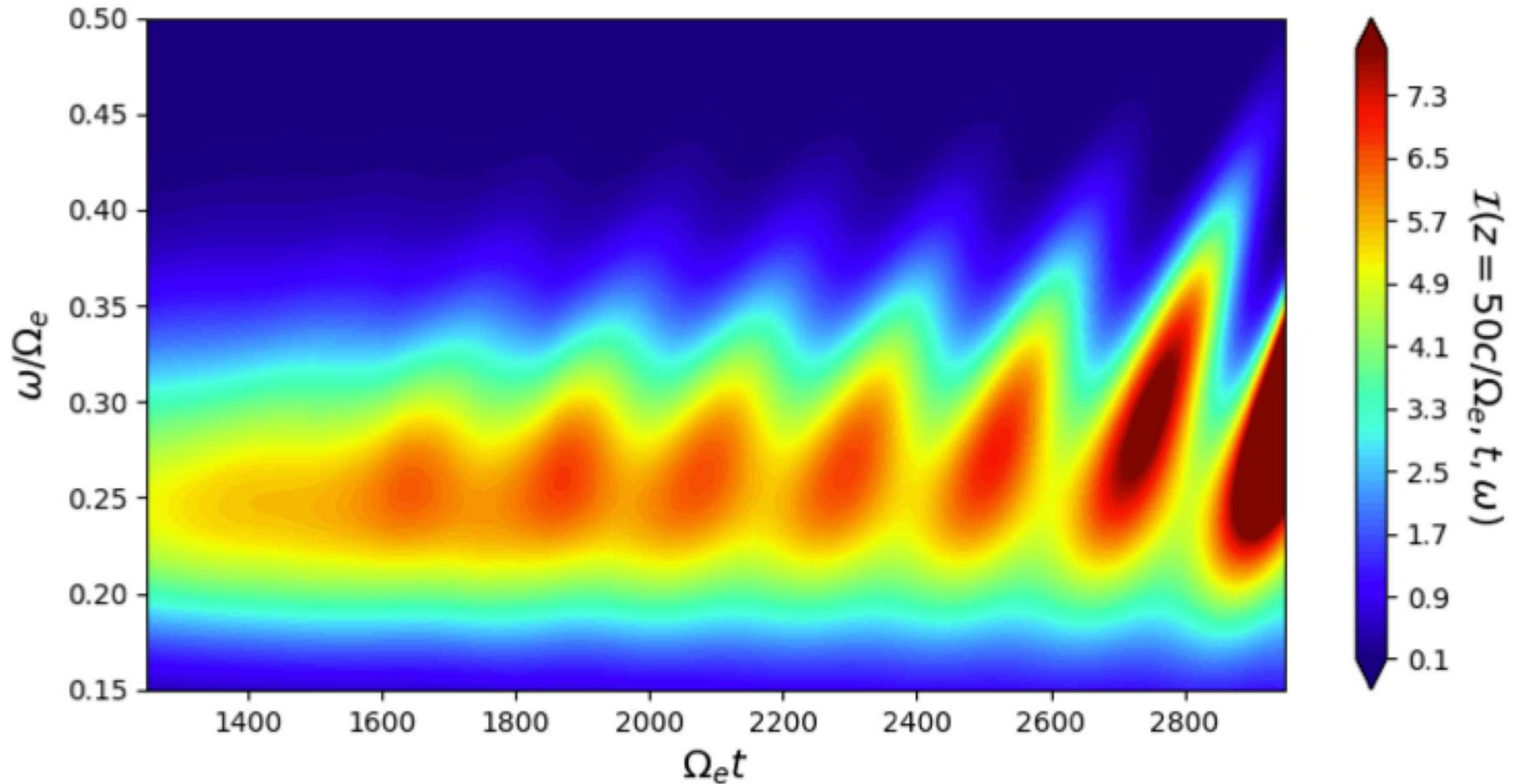
- Analytic result derived for the first time
- Self similar solution that ballistically propagates

$$\frac{\partial^2}{\partial t^2} \bar{\Gamma}_{NL}(\omega) \simeq \left(\sum_{k'} \frac{\langle \langle \omega_{trk'}^4 \rangle \rangle}{4 (1 - v_{r\omega'}/v_{g\omega'})^4} \right) \frac{\partial^2}{\partial \omega^2} \bar{\Gamma}_{NL}(\omega)$$

$$\frac{\partial \omega}{\partial t} = \pm \frac{1}{2} \frac{\langle \langle \omega_{trk}^4 \rangle \rangle^{1/2}}{(1 - v_{r\omega}/v_{g\omega})^2}$$

Numerical results

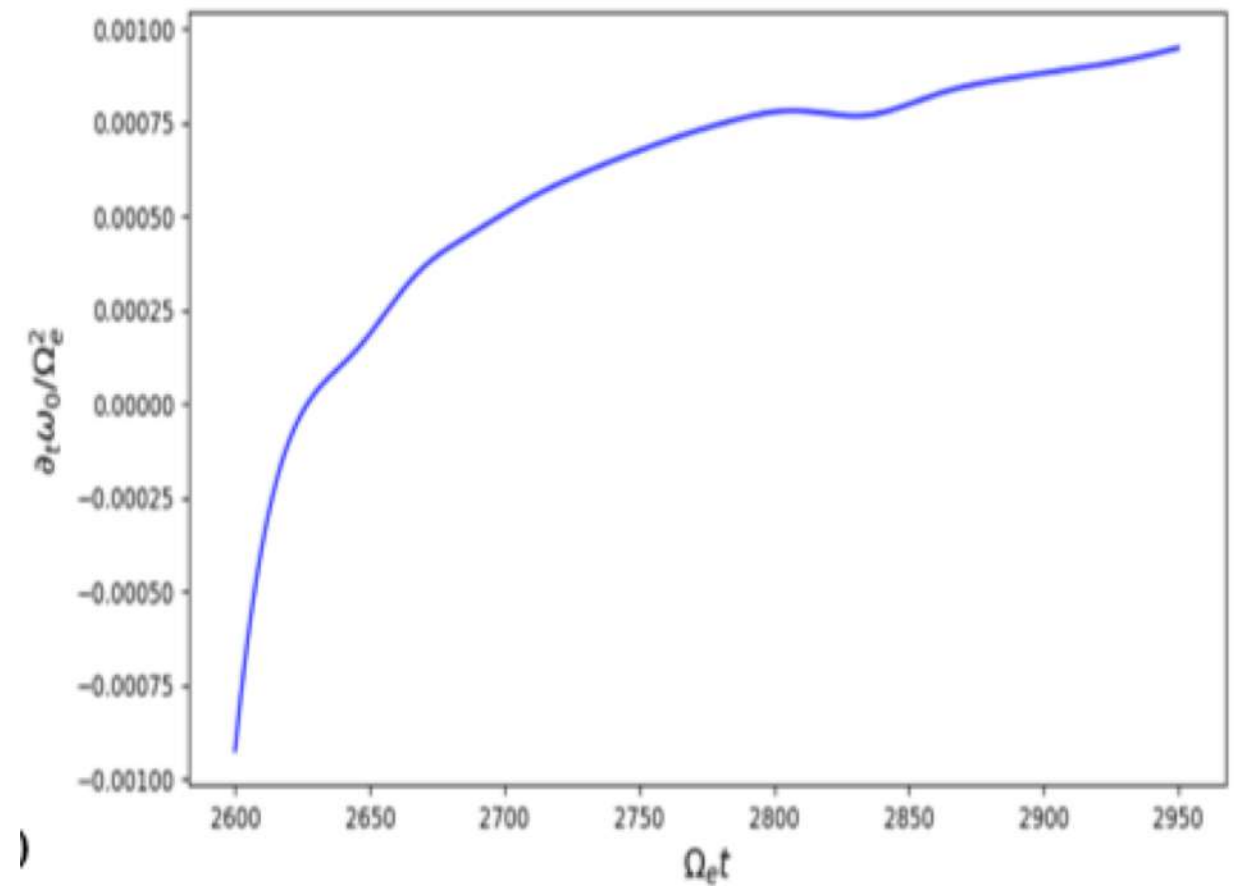
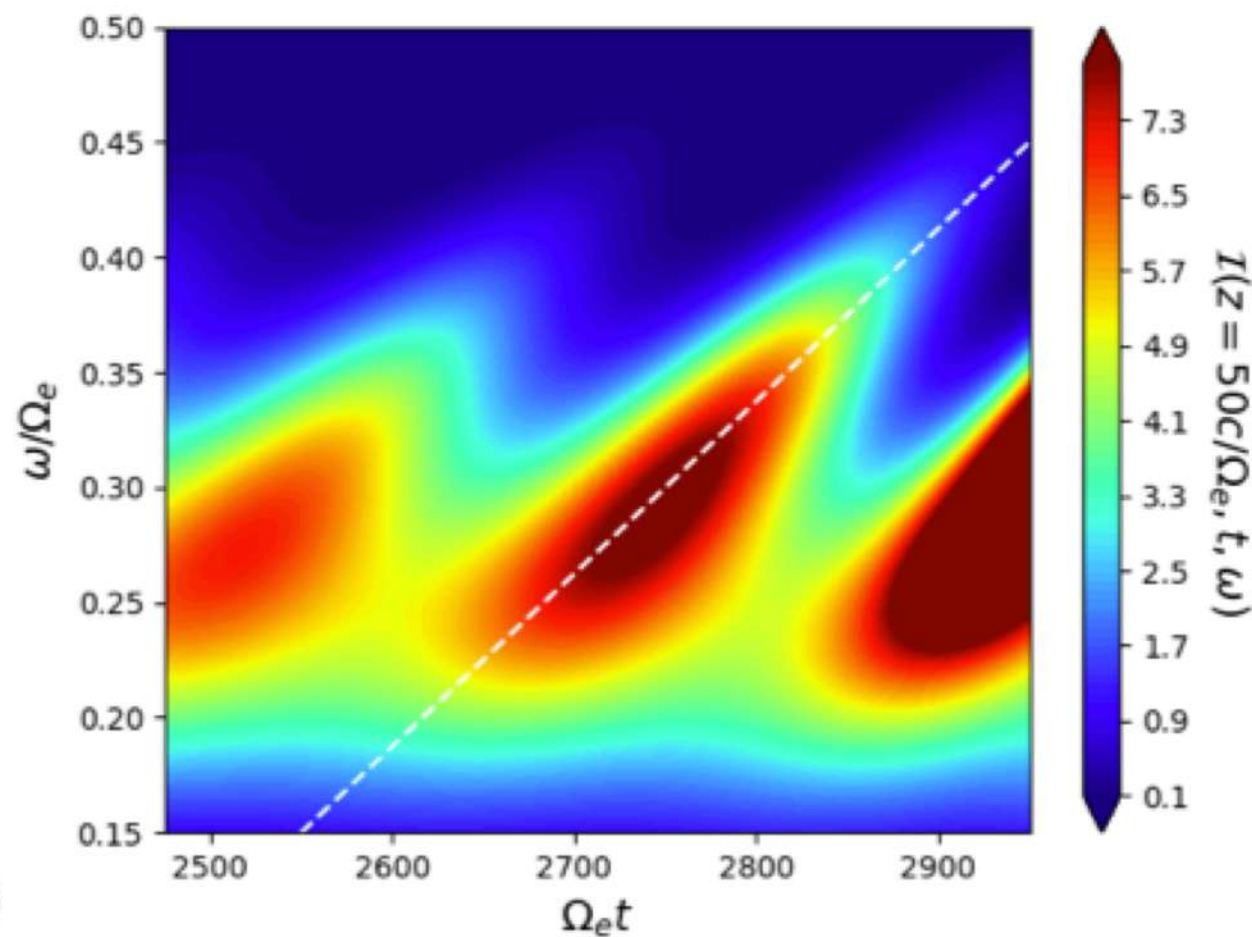
- Reduced Dyson description:



Numerical results

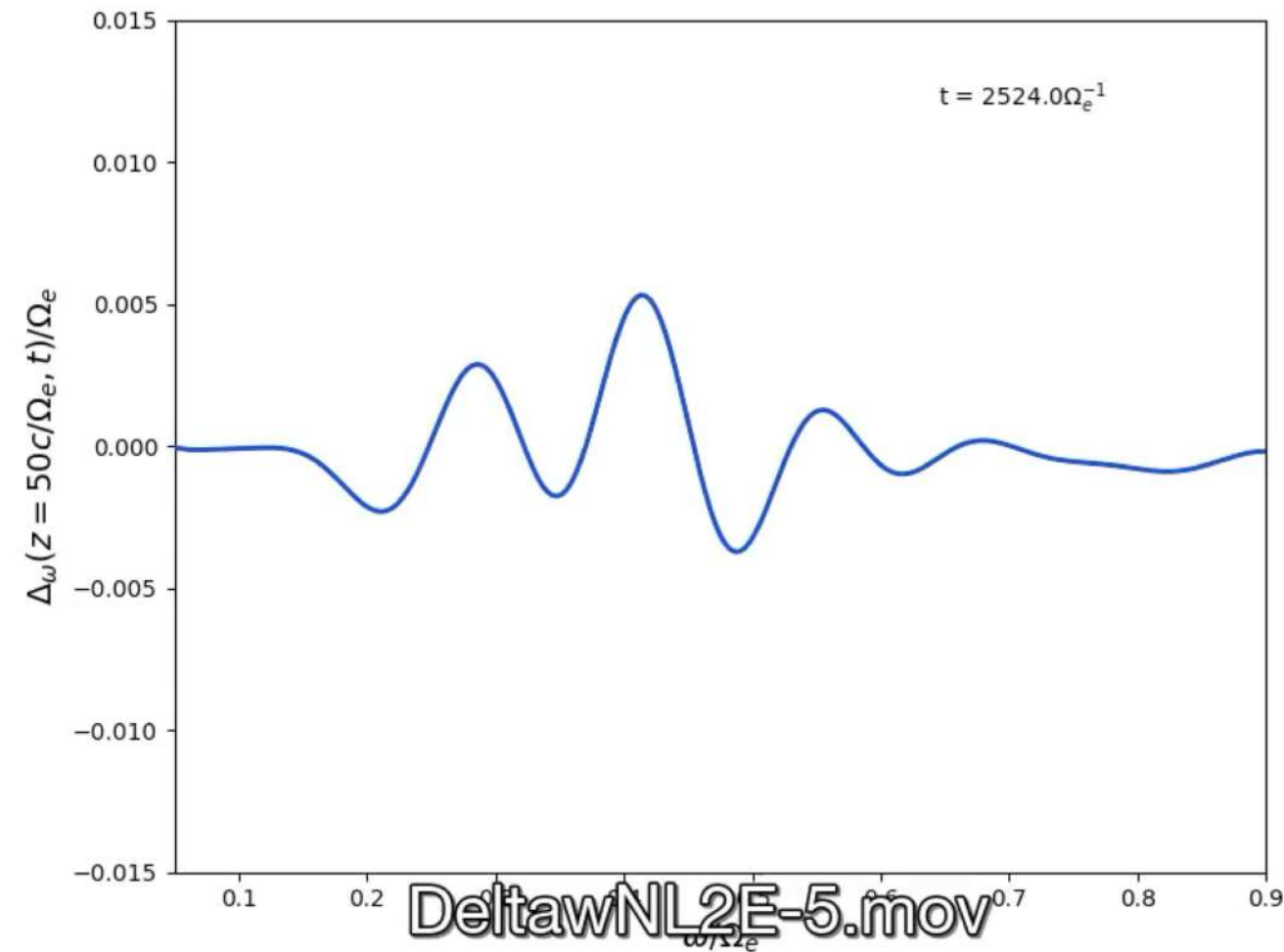
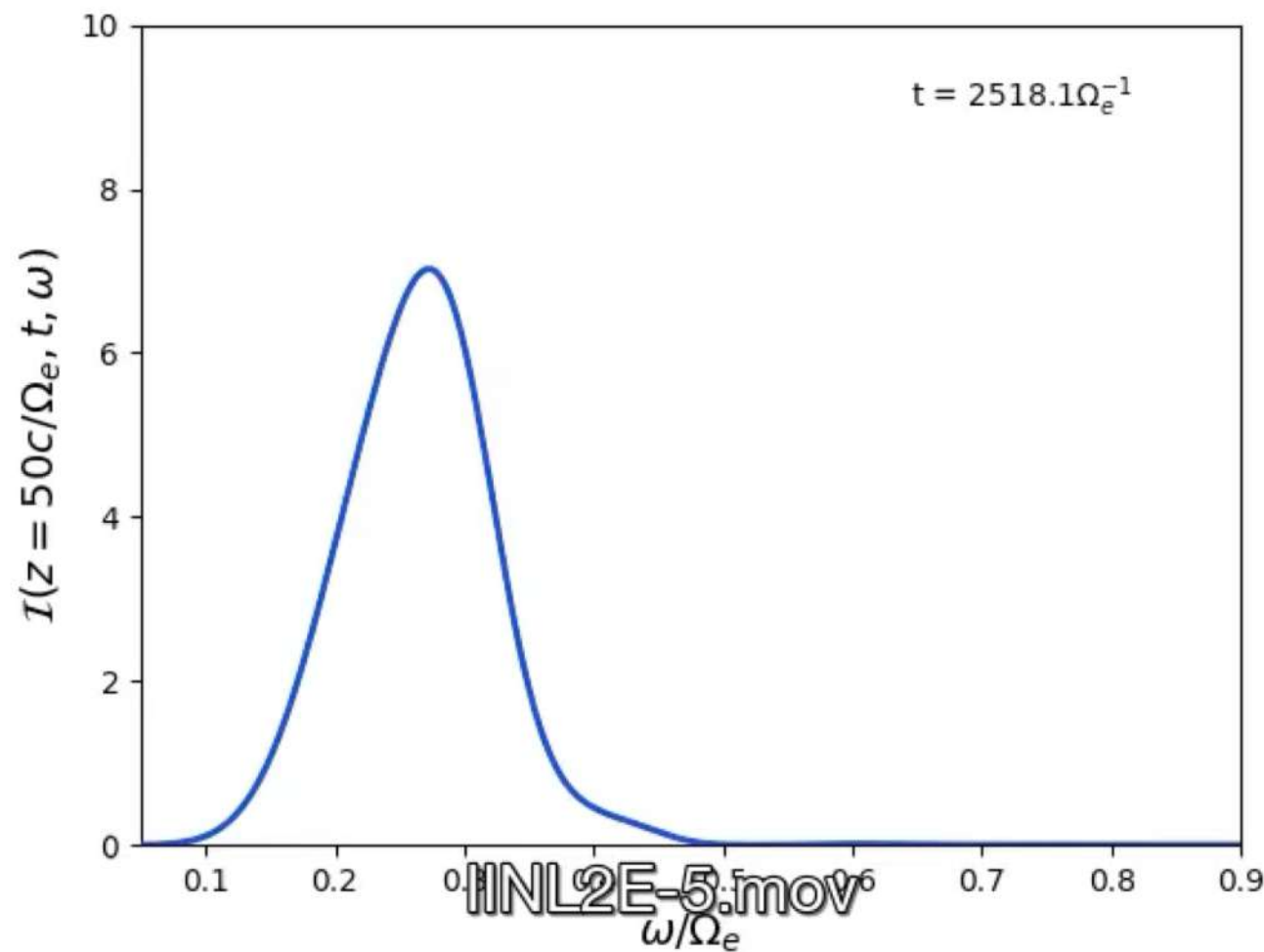
Reduced Dyson description

- Chirping rate 7.5×10^{-4} consistent with PIC simulation



A one - chirp show

- A one – chirp show: single chorus element



Universal chirping phenomena

L. Chen and F. Zonca, [RMP 88, 015008 \(2016\)](#)

- Wide variety of frequency sweeping phenomena
- **Universal process** when:
 - Spontaneous emission (instability)
 - Non-perturbative (dispersiveness, continuum)
 - Resulting in **secular propagation of phase space resonance structures** at rate $\sim$ amplitude
 - E.g., chorus chirping (Vomvoridis, 1982; ...)

$$\frac{\partial \omega}{\partial t} = \pm \frac{1}{2} \frac{\langle \langle \omega_{trk}^4 \rangle \rangle^{1/2}}{(1 - v_{rw}/v_{gw})^2}$$

• arXiv:2107.03151; [PDF](#)

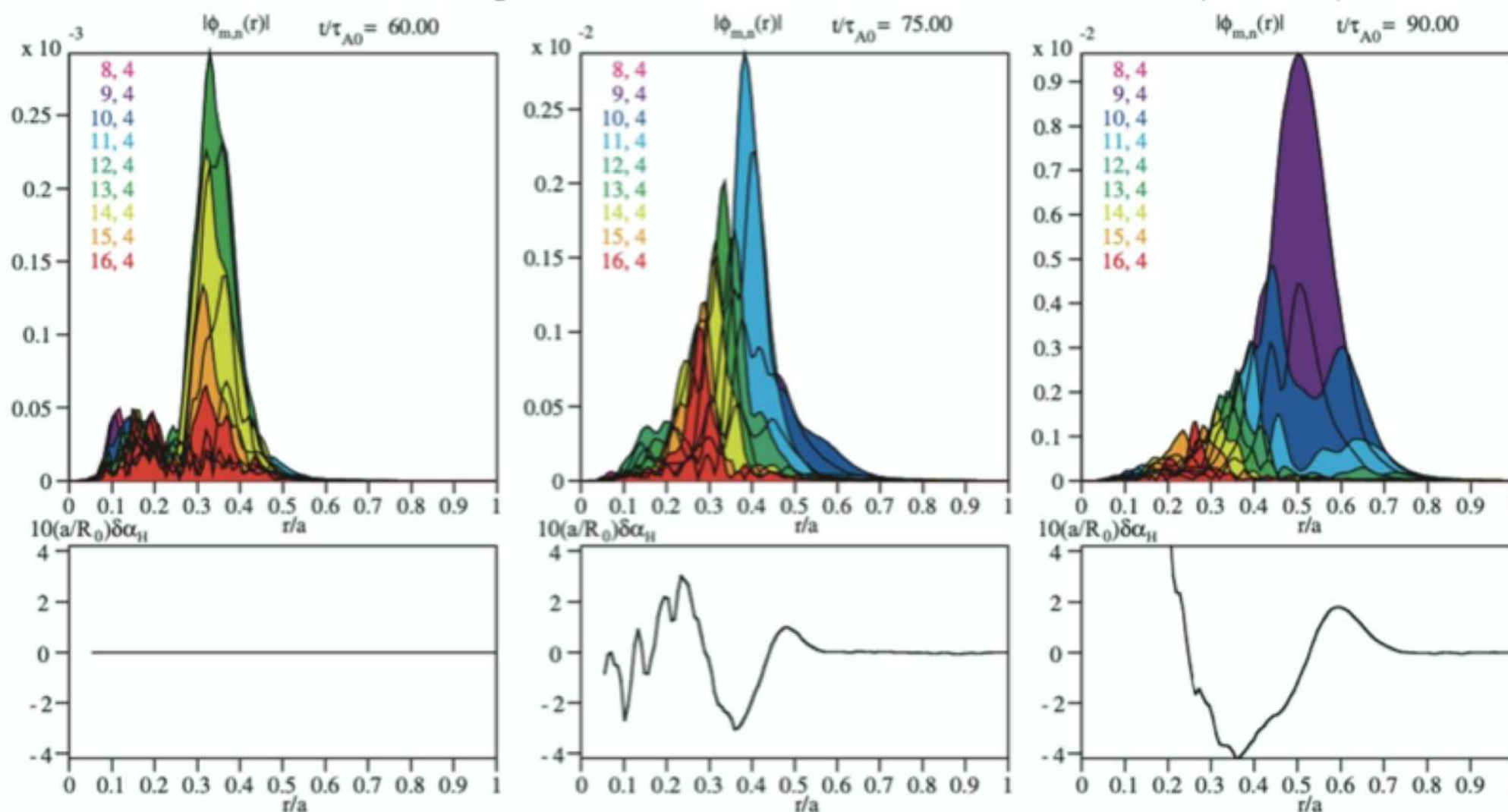
• arXiv:2106.15026 [PDF](#)

- Predicted and observed in fusion plasmas:
(EPM, fishbones)

Back to fusion plasmas - EPM

Theoretical framework systematically verified by NL Hybrid MHD GK simulations (HMGC): (Xin Wang seminar Apr. 09/2021; Sergio Briguglio seminar Apr. 23/2021; **Chirping explained**)

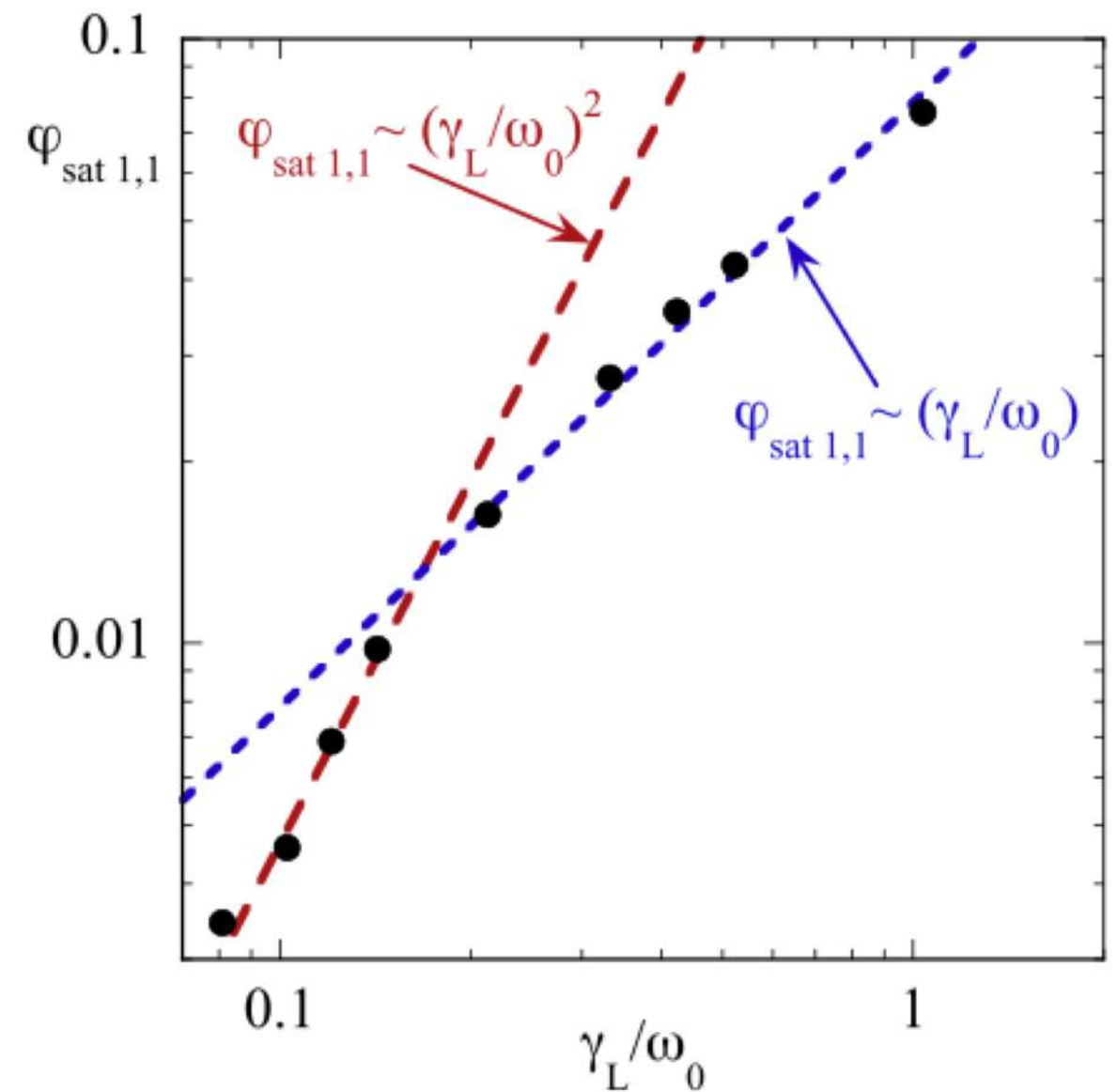
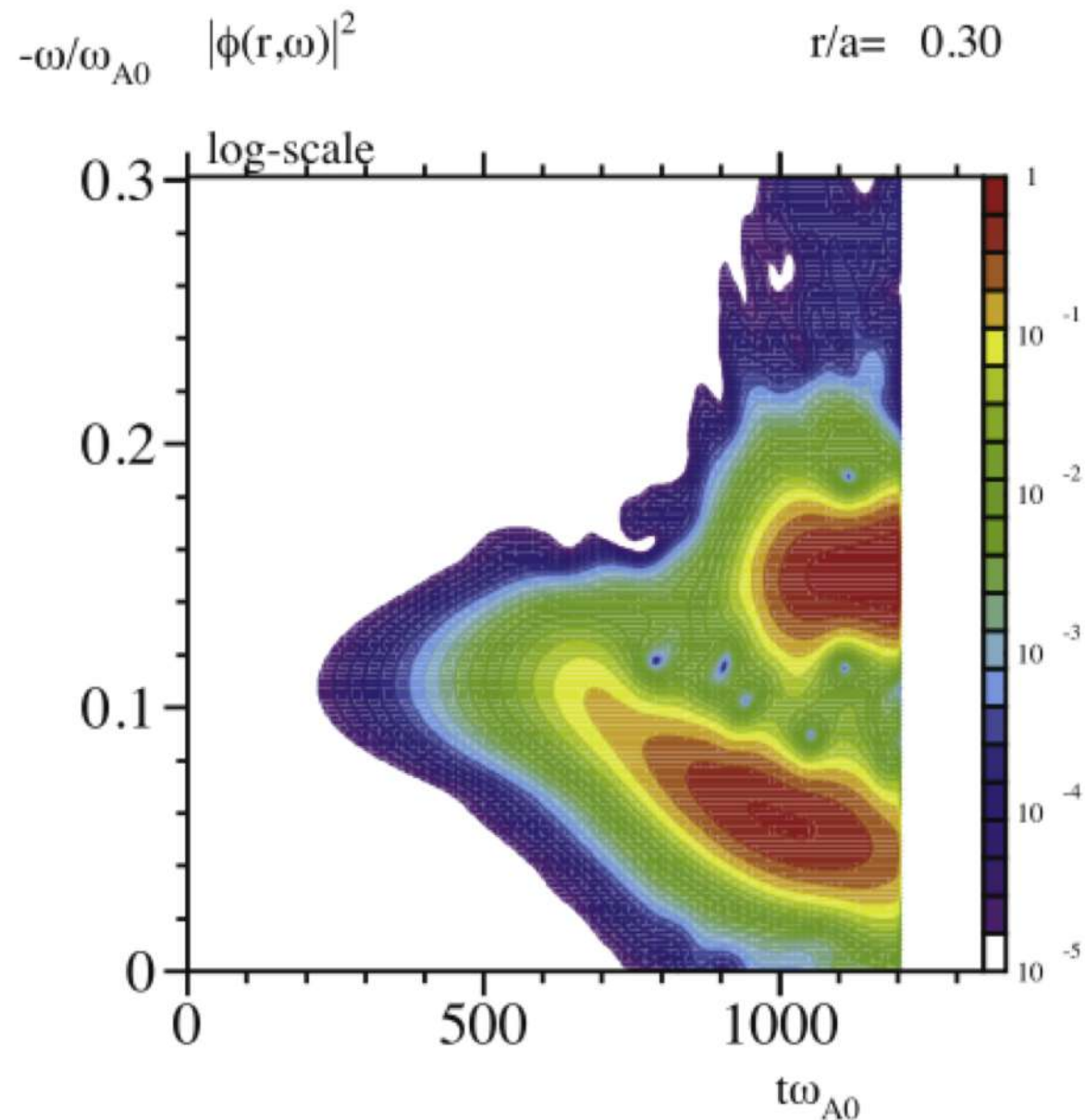
Evidence of EP avalanche [Z. et al Nucl. Fusion 45 (2005) 477 – 484]



Back to fusion plasmas - fishbones

Theoretical framework systematically verified by NL Hybrid MHD GK simulations (HMGC&HYMAGYC):

(G. Vlad et al. New J. Phys. 18 (2016) 105004)



**XHMGC e-fishbones;
weakly hollow q profile**

A hierarchy of models

Connecting Chorus with DSM:

- The DSM is made of two components: the Dyson Schwinger Equation (DSE) and NLSE
- The NLSE evolves the spectrum consistent with the nonlinear equilibrium solution; the DSE evolves the nonlinear equilibrium given the spectrum
- The chorus problem is the analogue of DSM, where the NLSE is replaced by the WKE

Recovering Quasilinear from DSM:

- This correspondence of DSM to chorus problem allows to illustrate how Quasilinear description is recovered in the proper limit

Dyson – like equation (DSE)

$$(\partial_t + v_{\parallel} \partial_z) f_0 = \frac{e}{m} \frac{v_{\perp}}{4} \text{Re} \sum_k i \left\{ \delta \bar{E}_k \left[\bar{\partial}_{\varepsilon k} + \frac{1}{v_{\perp}^2} \right. \right. \\ \left. \left. \times \left(1 - \frac{kv_{\parallel}}{\omega_k} \right) \right] \delta \bar{f}_k^* - \delta \bar{E}_k^* \left[\bar{\partial}_{\varepsilon k} + \frac{1}{v_{\perp}^2} \left(1 - \frac{kv_{\parallel}}{\omega_k} \right) \right] \delta \bar{f}_k \right\}$$

Closed by ...

$$\delta \bar{f}_k = \frac{e}{m} \mathcal{P}_k v_{\perp} \delta \bar{E}_k \bar{\partial}_{\varepsilon k} f_0 ,$$

$$\mathcal{P}_k \equiv [\Omega + kv_{\parallel} - \omega_k - i(\partial_t + v_{\parallel} \partial_z) - \Delta_k]^{-1}$$

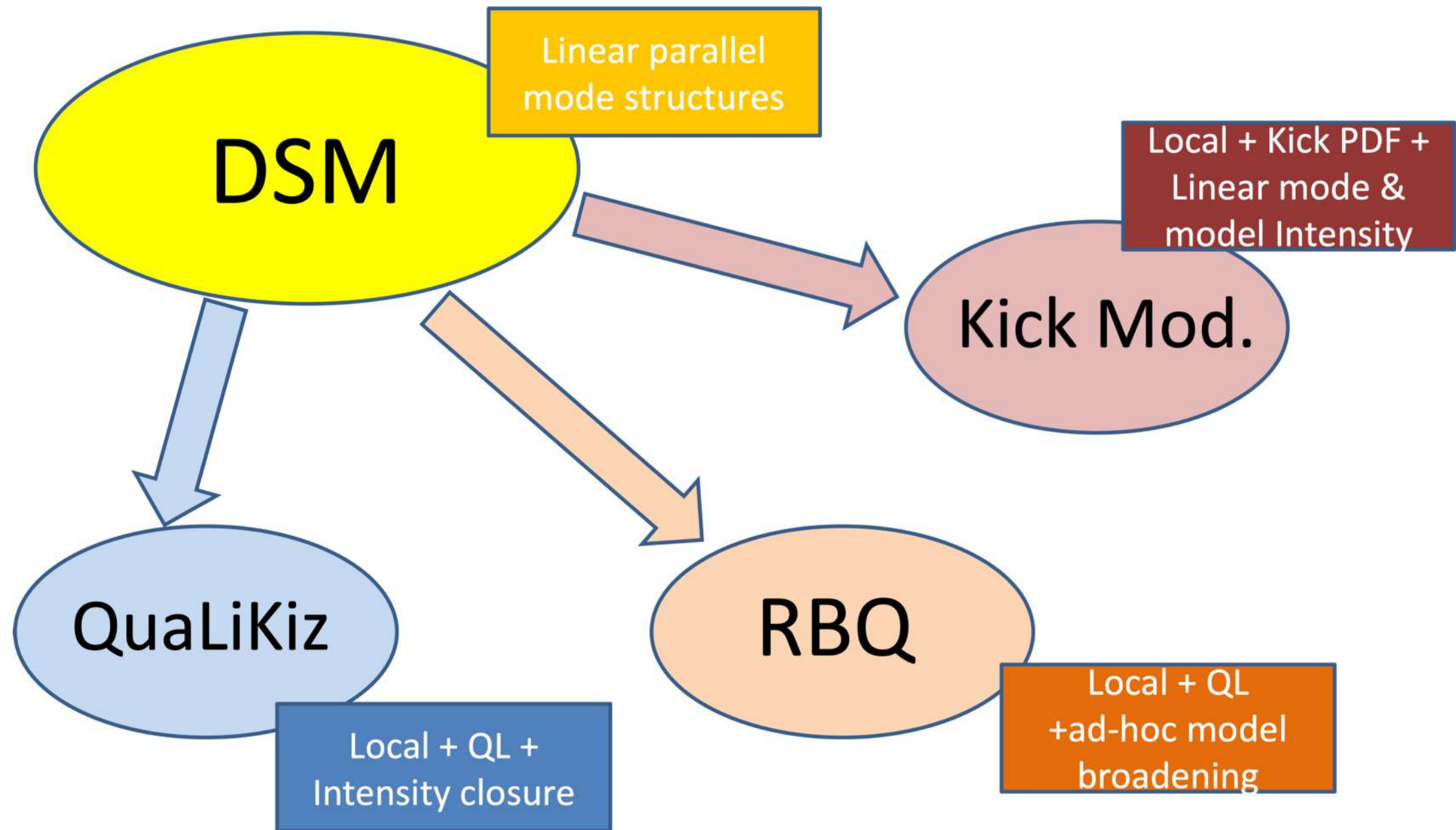
resonance broadening

slow spatiotemporal dep.

1. Propagator becomes the linear one
2. The spectrum is considered broad (w.r.t. ω_{trap}) small Kubo #
3. Resonances overlap and $\omega_{\text{trap}} \gg 1/\tau_{\text{diff}} \sim 1/\tau_{\text{NL}}$
4. Near marginal stability: γ even smaller than $1/\tau_{\text{diff}} \sim 1/\tau_{\text{NL}}$

Chorus and fusion plasmas - II

Direct connection with Dyson Schrödinger Model (DSM) :
(ZCFQ seminar Nov 20/2020)



Conclusions

- The theoretical framework for chorus emission is fully developed based on the DSE+WKE solution. Supported by extensive numerical verification.
- It describes universal chirping phenomena in space/lab plasmas
- The chorus problem is the analogue of DSM, where the NLSE is replaced by the WKE. This provides a direct route to demonstration that the DSE+WKE/NLSE solution reduces to QL in the proper limit
- This illustrates the DSM as a robust framework for a hierarchy of reduced models, yielding QL and local CG
- These results suggest the systematic application of DSM to a broader class of problems dealing with transport in phase space (beyond AE/EPM/fishbones) → ATEP/TSVV#10 EUROfusion Proj.

THANK YOU!

Your questions are welcome