

Resonant axisymmetric modes in Tokamak plasmas

Franco Porcelli

Polytechnic University of Turin, Italy

In collaboration with:

A. Yolbarsop, U. of Science and Technology of China

T. Barberis, Polytechnic U. of Turin

R. Fitzpatrick, IFS, U. of Texas at Austin, USA



**POLITECNICO
DI TORINO**

OUTLINE

INTRO – X-point resonance, vertical displacements, feedback stabilization.

PART I – The math of X-point singularity, and (perhaps) a surprising result.

PART II – A new fast ion instability?

CONCLUDING REMARKS – A new line of research.

INTRODUCTION

Any unexplored concepts in
tokamak plasma theory out
there?



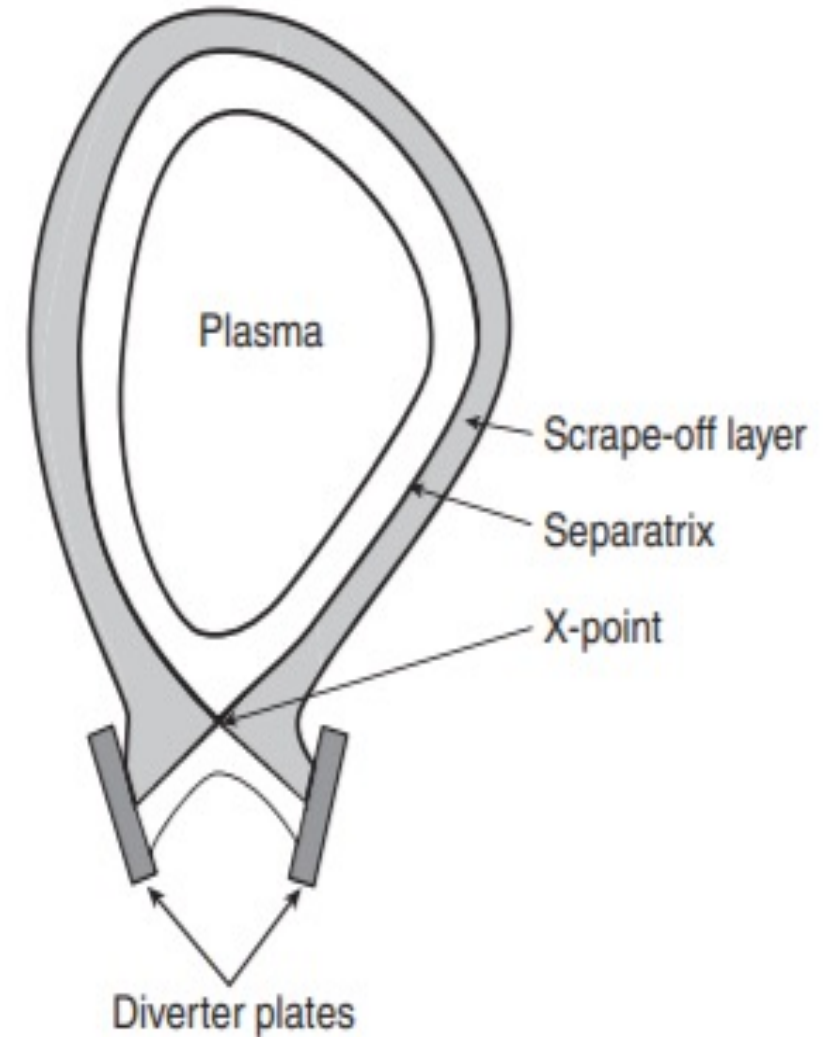
X-point

$$\mathbf{B} \cdot \nabla \xi = i(n/R)B_T \xi + i B_p(\psi, \vartheta) \sum_m (m/r) \xi_m = 0$$

Resonance condition at the X-point
satisfied for $n=0$ and any m

because $B_p(\psi, \vartheta)=0$ on the X-point.

The toroidal field line going through the
magnetic X-point is resonant to $n=0$
perturbations!



Schematic diagram of a divertor
(single null configuration)

Axisymmetric Modes

Axisymmetric ($n=0$) modes in shaped toroidal plasmas are global plasma modes, normally associated with VDEs.

Because of the X-point resonance, **current sheets localized along the magnetic separatrix are likely to form**. This may have an important impact on plasma edge activity.

Ref.: A. Yolbarsop, F. Porcelli, R. Fitzpatrick, Nuclear Fusion Letters 2021.

Furthermore, a stable **$n=0$ mode tends to acquire an Alfvénic oscillation frequency**. As a result, a **“resonant” interaction with energetic particles** can lead to a new type of fast ion instability.

Ref.: T. Barberis, F. Porcelli and A. Yolbarsop, submitted for publication (Oct. 2021)



But then, why nobody has worried about the $n=0$ resonance?

I am not sure...

Perhaps... because these modes are taken care of by the feedback stabilization system?

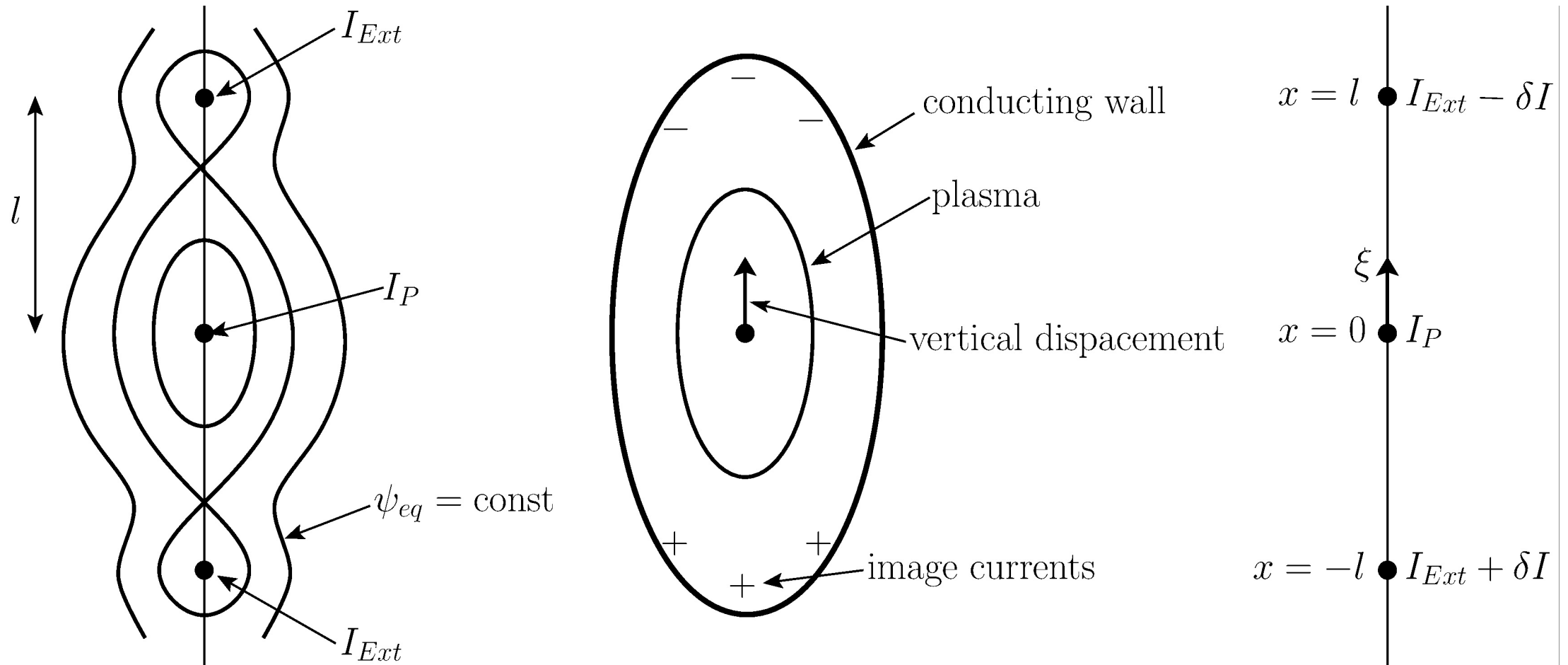
Let us review the philosophy behind passive and active feedback stabilization of vertical plasma displacements.

A heuristic model:
three parallel current wires.

Lister et al, NF 1990;
Lazarus et al, NF 1990.



Three-current-wires model



Cubic dispersion relation

$$\mu \ddot{\xi} + \frac{1}{\tau_R} \dot{\xi} + \omega_0^2 \xi - \frac{1}{\tau_R} \frac{\omega_0^2}{D-1} = 0$$

where

$$\omega_0^2 = (D-1)e_0\omega_{pA}^2; \quad D = \frac{b^2+a^2}{(b-a)^2} \frac{b_w-a_w}{b_w} \quad (\text{Laval et al, PF1974})$$

$D > 1$ for passive feedback stabilization

Three roots:

$$\omega = \pm \omega_0 - \frac{iD}{2(D-1)\tau_R} \quad \text{stable, oscillatory}$$

$$\omega = i\gamma = i \frac{1}{(D-1)\tau_R} \quad \text{unstable (requires active feedback)}$$

Note: only two oscillatory roots for $\tau_R \rightarrow \infty$:

$$\omega^2 = \omega_0^2 \quad (\text{ideal wall limit})$$

Comments

- Only the weakly growing, zero frequency mode is considered for active feedback stabilization.
- **No X-point resonance effects.**
- **Hey! What about the weakly damped, high frequency solutions?**

PART I

The math of X-point singularity, and (perhaps) a surprising result.

For details, see:

IOP Publishing | International Atomic Energy Agency

Nuclear Fusion

Nucl. Fusion 61 (2021) 114003 (5pp)

<https://doi.org/10.1088/1741-4326/ac27c5>

Letter

Impact of magnetic X-points on the vertical stability of tokamak plasmas

A. Volbarsop^{1,2}, F. Porcelli^{1,*} and R. Fitzpatrick³

¹ DISAT, Polytechnic University of Turin, Torino 10129, Italy

² KTX Laboratory, School of Nuclear Science and Technology, University of Science and Technology of China, Hefei, 230022, China

³ Institute for Fusion Studies, University of Texas at Austin, United States of America

Reduced ideal-MHD model

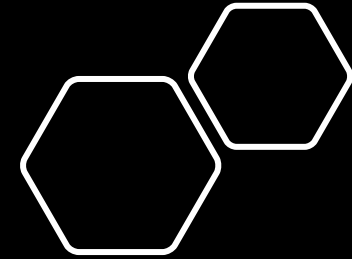
$$\mathbf{B} = \mathbf{e}_z \times \nabla \psi + B_z \mathbf{e}_z$$

$$\mathbf{v} = \mathbf{e}_z \times \nabla \varphi.$$

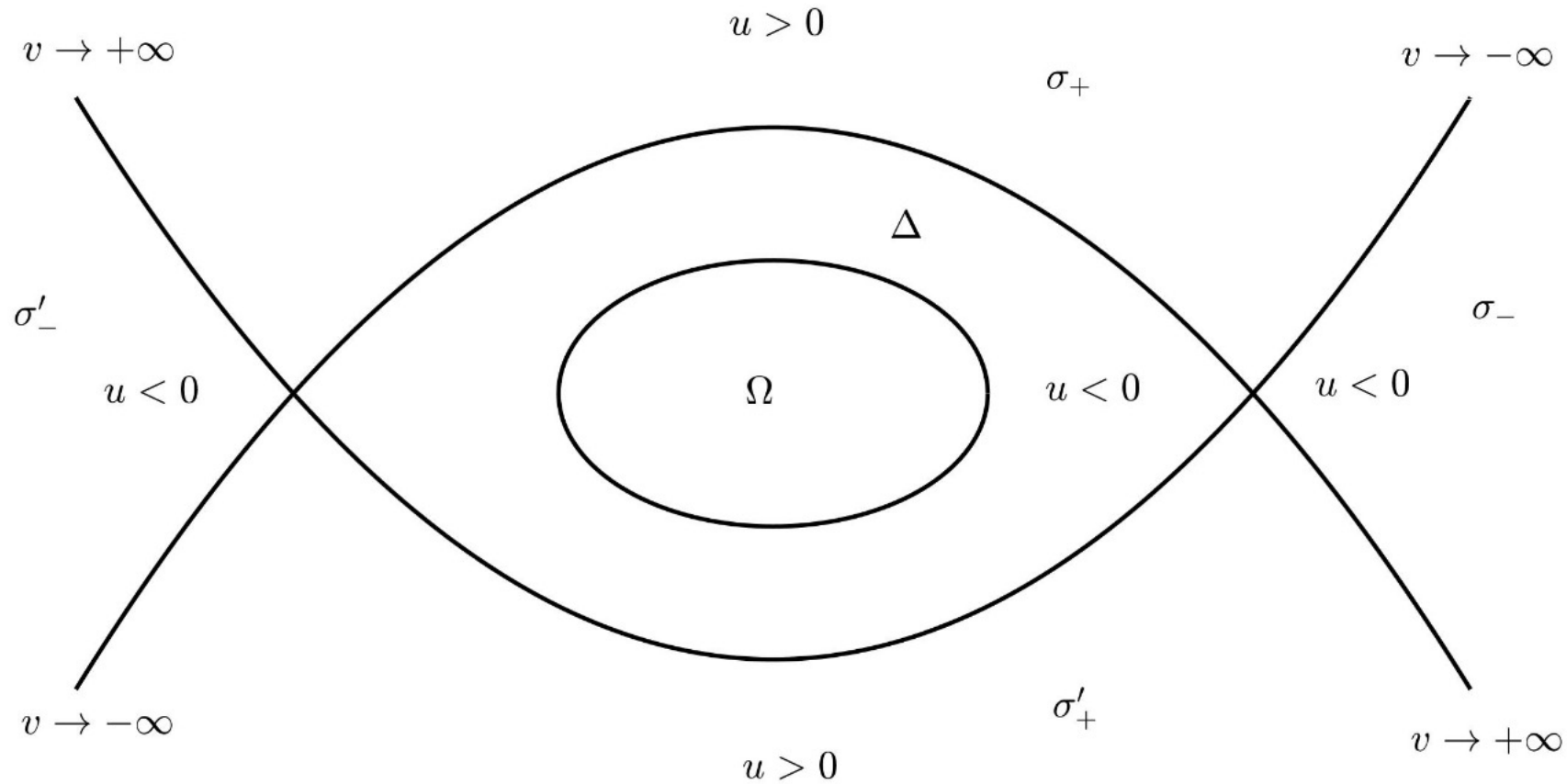
$$\frac{\partial \psi}{\partial t} + [\varphi, \psi] = 0$$

$$\frac{\partial}{\partial t} \nabla \cdot (\varrho \nabla \varphi) + [\varphi, U] = [\psi, J]$$

$$[\chi, \eta] = \mathbf{e}_z \cdot \nabla \chi \times \nabla \eta, \quad J = \nabla^2 \psi, \quad U = \nabla^2 \varphi.$$



Equilibrium magnetic structure



$$J_{eq} = J_0 H(\psi_b - \psi_{eq})$$

$$\varrho_{eq} = \varrho_0 H(\psi_X - \psi_{eq})$$

Plasma occupies Ω and Δ regions

Equilibrium

Uniform current density within
an elliptical boundary

$$J_{eq} = J_0 H(\mu_b - \mu) = J_{eq}(\psi_{eq})$$

Elliptical coordinates

$$x = A \cosh \mu \cos \theta$$

$$y = A \sinh \mu \sin \theta$$

$$A = \sqrt{b^2 - a^2}$$

$$a = A \sinh \mu_b$$

$$b = A \cosh \mu_b$$

Equilibrium solution

$$[\psi_{eq}, J_{eq}] = 0 \quad \Rightarrow \quad \nabla^2 \psi_{eq} = J_0 H(\mu_b - \mu)$$

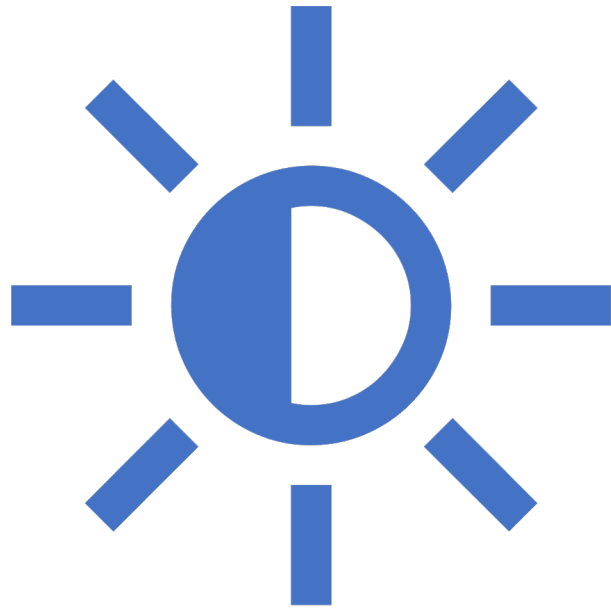
Solution:

$$\psi_{eq}^-(x, y) = \frac{1}{2} \left(\frac{x^2}{b^2} + \frac{y^2}{a^2} \right) \quad \Omega \text{ region}$$

$$\psi_{eq}^+(\mu, \theta) = \frac{1}{2} + \alpha^2 \left\{ \mu - \mu_b - \frac{e_0}{2} \sinh [2(\mu - \mu_b)] \cos(2\theta) \right\} \quad \Delta \text{ and } \sigma \text{ regions}$$

$$\nabla^2 \chi = h^{-2} (\partial^2 \chi / \partial \mu^2 + \partial^2 \chi / \partial \theta^2)$$
$$h = 1/|\nabla \mu| = 1/|\nabla \theta|$$

Refs: Porcelli&Yolbarsop, PoP 2019; Gajewski, PF 1972.



Perturbations:

The ideal-MHD constraint
at the magnetic X-point

and

the rigid-shift solution.

EXACT ANALYTIC SOLUTION

Rigid-vertical-shift stream function

$$\varphi = -\gamma \xi y$$

Elliptical coordinates

$$x = A \cosh \mu \cos \theta$$

$$A = (b^2 - a^2)^{1/2}$$

$$y = A \sinh \mu \sin \theta$$

X- point coordinates

$$\mu = \mu_X = 2\mu_b \quad \text{and} \quad \theta = \theta_X = 0, \pi$$

Stream function in elliptical coordinates:

$$\varphi(\mu, \theta) = -\gamma A \sinh \mu \sin \theta$$

Use flux-freezing condition to obtain perturbed flux

$$\gamma \tilde{\psi} + [\tilde{\varphi}, \psi_{eq}] = 0,$$

$$\tilde{\psi}_\Omega(\mu, \theta) = -\xi_0 \frac{\cosh(\mu)}{\cosh \mu_b} \cos \theta.$$

$$\tilde{\psi}_\Delta(\mu, \theta) = \xi_0 \frac{\sinh(\mu - 2\mu_b)}{\sinh \mu_b} \cos \theta.$$

Note:
single m=1
harmonic in
elliptical angle;

perturbed flux
vanishes at
X-points.

**Vacuum solution
(no wall)**

$$\tilde{\psi}_{\sigma\pm}(\mu, \theta) = \sum c_m \exp(-\tilde{m}\mu) \cos(m\theta).$$

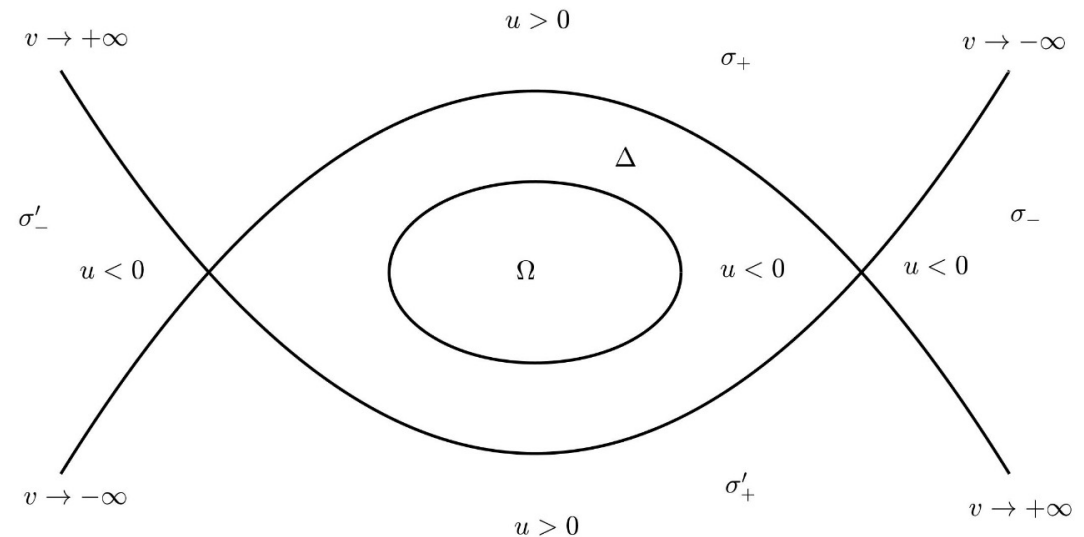
However, the separatrix is not a $\mu = \text{const}$ magnetic surface, and so analytic determination of coefficients c_m is an inherently intractable problem.

Analytic progress is possible if **flux coordinates** (u, v) are used:

$$u = \alpha^{-2}(\psi_{eq} - \psi_X) = \mu - 2\mu_b + \frac{e_0}{2} \{ \sinh(2\mu_b) - \sinh[2(\mu - \mu_b)] \cos(2\theta) \}$$

$$v = \theta - \frac{e_0}{2} \cosh[2(\mu - \mu_b)] \sin(2\theta)$$

$$e_0 = \frac{b^2 - a^2}{b^2 + a^2},$$

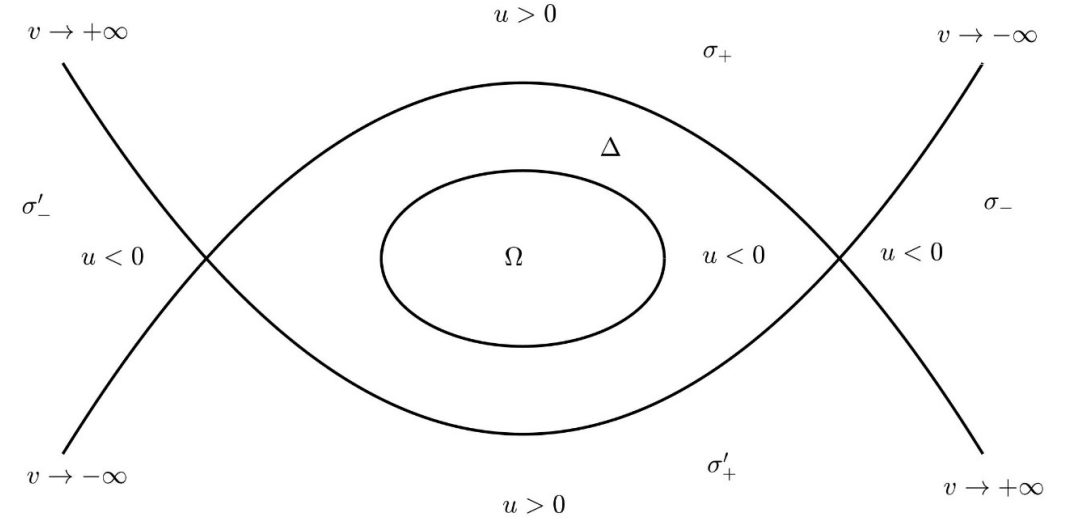


Since $\nabla^2 \tilde{\psi}_\Delta = 0$

$$\tilde{\psi}_\Delta(u, v) = \xi_0 \sum_{m, \text{odd}}^{\infty} [\alpha_m \cosh(mu) + \gamma_m \sinh(mu)] \cos(mv).$$

$\tilde{\psi}_\Delta(u, v) = \tilde{\psi}_\Delta(\mu, \theta)$. We can use this to determine the α_m, γ_m coefficients.

$$\tilde{\psi}_\Delta(\mu, \theta) = \xi_0 \frac{\sinh(\mu - 2\mu_b)}{\sinh \mu_b} \cos \theta.$$



After straightforward algebra, we obtain $\alpha_m = a_m + b_m$, $\gamma_m = -a_m + b_m$

$$a_m = \frac{e^{mu_b}}{2m} \left[\left(\frac{b}{a} + 1 \right) J_{\frac{m-1}{2}} \left(\frac{me_0}{2} \right) + \left(\frac{b}{a} - 1 \right) J_{\frac{m+1}{2}} \left(\frac{me_0}{2} \right) \right]$$

$$b_m = \frac{e^{-mu_b}}{2m} \left[\left(\frac{b}{a} - 1 \right) J_{\frac{m-1}{2}} \left(\frac{me_0}{2} \right) + \left(\frac{b}{a} + 1 \right) J_{\frac{m+1}{2}} \left(\frac{me_0}{2} \right) \right]$$



Mathematics of X-point singularity: non-analiticity

Consider magnetic flux on separatrix

$$\psi_{\Delta}(0, v) = \sum_{m \text{ odd}} \alpha_m \cos mv$$

Fourier spectrum at large m

$$\alpha_m \sim p e_0^{1/2} / m^{3/2} \rightarrow \psi_{\Delta}(0, v) \text{ NOT an analytic function of } v$$

Indeed:

$$\psi_{\Delta}(0, v) \sim \sqrt{\frac{\pi}{2}} p e_0^{1/2} |v|^{1/2} \quad \text{as } v \rightarrow 0$$

$$p \approx (2e_0/\pi)^{1/2} \text{ for small } e_0$$

Continuity along the separatrix determines the perturbed flux in the vacuum region. This involves a parameter q .

Along the separatrix:

$$\tilde{\psi}_{\sigma^+}(0, \nu) = \int_0^\infty dt \alpha(t) \sin \left[\left(\frac{\pi}{2} - \nu \right) t \right] = \tilde{\psi}_{even}(0, \nu) + \tilde{\psi}_{odd}(0, \nu)$$

even: $q = p/2$

odd: $q = -p/2$

$$p \approx (2e_0/\pi)^{1/2} \text{ for small } e_0$$

Current sheet along the separatrix:

$$\tilde{J}(u, v) \sim j_X(v) \delta(u)$$

where

$$j_X(v) \sim e_0^{3/2} \left(\frac{p}{2} + q \right) \xi_0 |v|^{1/2} \quad \text{as } v \rightarrow 0$$

Growth rate:

$$\gamma^2 = -\sqrt{\frac{\pi a}{2b}} \frac{1}{\alpha^2} \left(q + \frac{p}{2} \right) e_0^{3/2} \omega_A^2$$

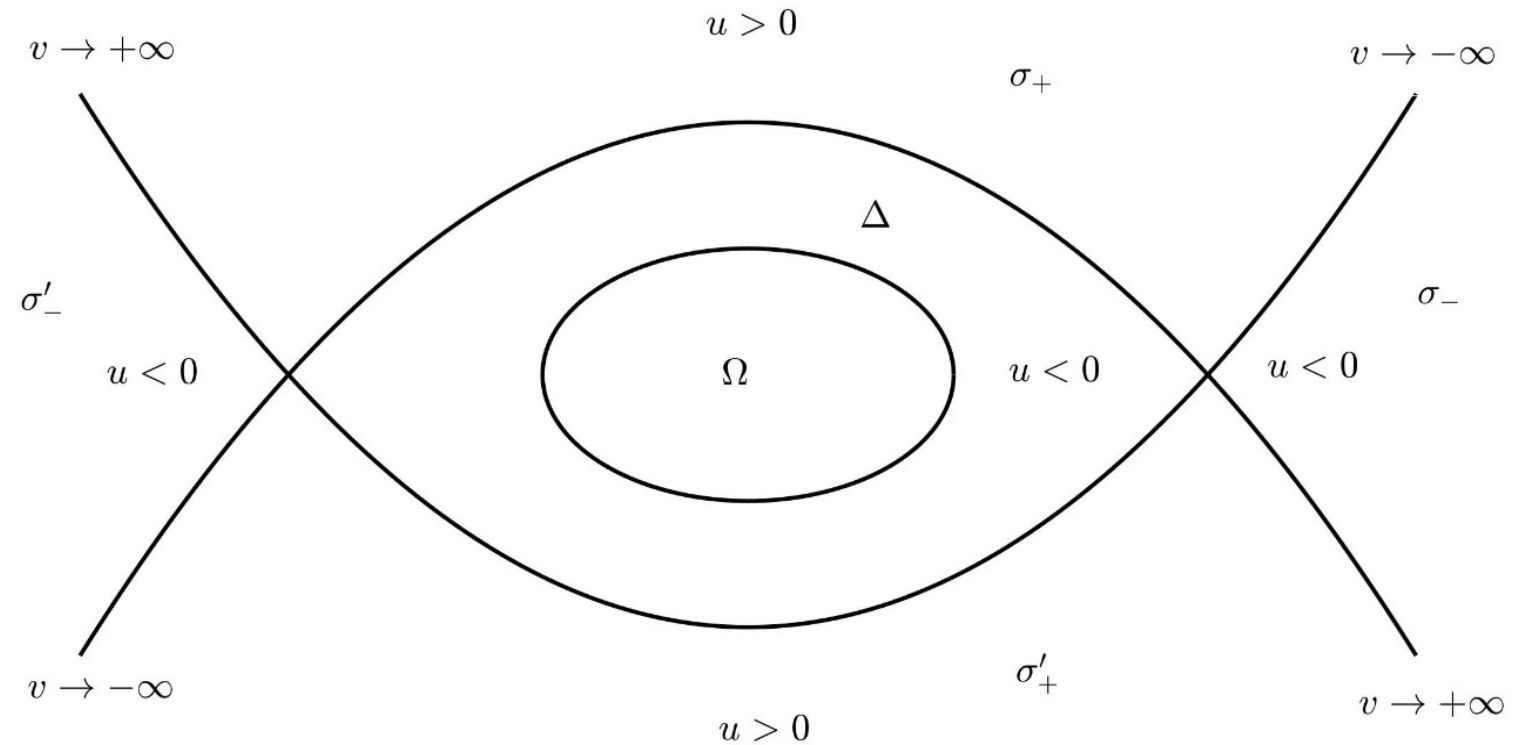
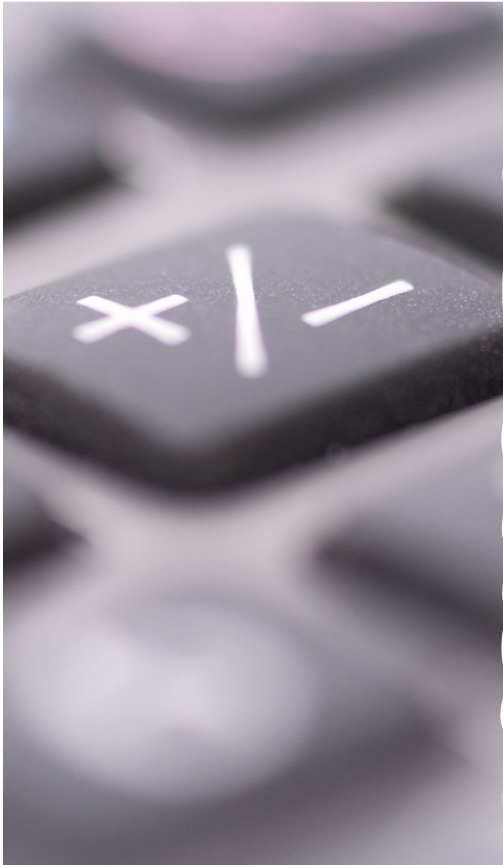
Two classes of modes:

even: $q=p/2$ - stable oscillatory roots - current sheet on separatrix
odd: $q=-p/2$ - neutrally stable with $\gamma = 0$ - no current sheet

$$p \approx (2e_0/\pi)^{1/2} \text{ for small } e_0$$

Mathematics of X-point singularity:

v non-periodic in vacuum region



Conclusions

The
mathematics
of the X-point
singularity is
established!

$n=0$ current sheets form along the magnetic separatrix, with $\tilde{j} \sim |\nu|^{1/2}$ as the X-point (at $\nu = 0$) is approached.

The surprising result is that, because of these current sheets, the ideal-MHD vertical displacement is stable even in the absence of a passive feedback stabilization wall.

Conclusions

Part I

The stabilization mechanism is a direct consequence of the ideal-MHD flux-freezing constraint on the X-points, which generates current sheets localized along the magnetic separatrix, exerting a force capable of pushing back the plasma in its vertical motion.

Analogies with the physics of current sheet formation from the nonlinear evolution of internal kink modes (*) and with the island coalescence problem (**) → the ideal-MHD constraint causes magnetic flux to pile up near the X-points, leading to perturbed localized currents and a stabilizing effect in the ideal-MHD limit.

(*) M. N. Rosenbluth, N. Y. Dagazian, and P. H. Rutherford, Phys. Fluids 16, 1894 (1973).

(**) J. M. Finn and P. K. Kaw, Phys. Fluids 20, 72 (1977).

Have these
axisymmetric
X-point
current
sheets been
observed?

Perhaps yes...

JET experiments:

- J. Lingertat et al., J. Nucl. Mat. 241, 402 (1997);
- E. R. Solano et al., Nucl. Fusion 57, 022021 (2016).

Numerical simulations (M3D-C¹; NIMROD; JOREK):

- I. Krebs, F. G. Artola, C. R. Sovinec, S. C. Jardin, K. J. Bunkers, M. Hoelzl, and N. M. Ferraro, Phys. Plasmas 10, 930 (2020).

PART II

A new fast ion instability?

One branch of the $n=0$ dispersion relation corresponds to oscillatory modes with a discrete frequency,

$$\omega \approx \pm e_0^{1/2} \omega_{pA}$$

not interacting with the Alfvén continuum, weakly damped by wall and/or plasma resistivity.

We find that this mode can be destabilized by the resonance with fast ions on both trapped and passing orbits

Experimental evidence

Observation of $n=0$ modes

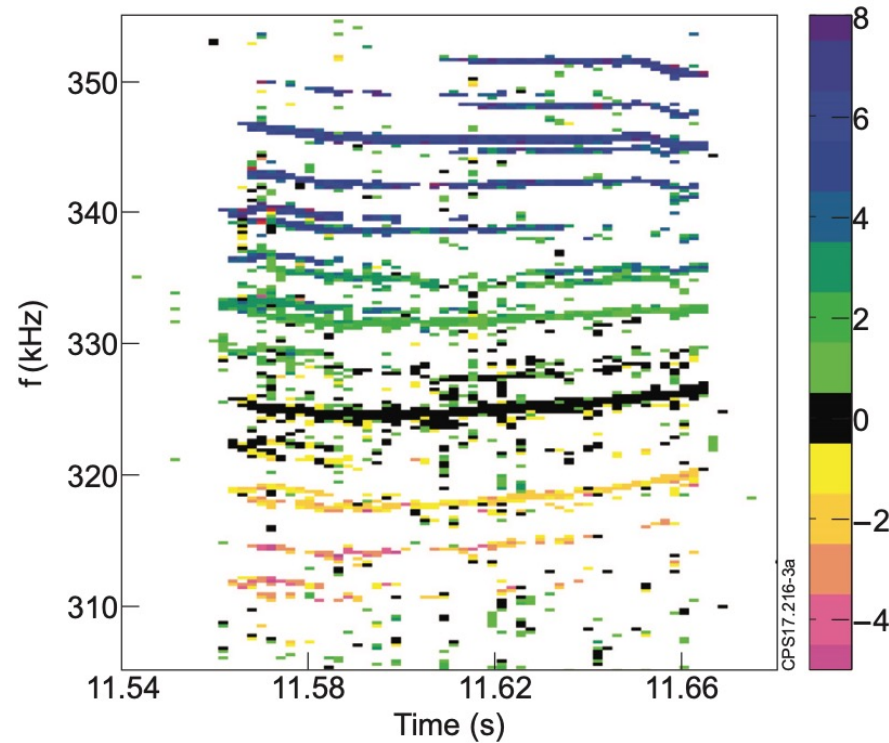


FIG. 3. Phase magnetic spectrogram showing the toroidal mode numbers of modes in the EAE frequency range excited in discharge #86775. Modes with toroidal mode numbers $-5 \leq n \leq 8$ were excited for a period of 0.1 s following a monster sawtooth crash at $t = 11.48$ s. $n = 0$ is shown in black.

Observation of saturated $n=0$ modes in $D-^3He$ plasmas at JET with NBI-ICRF.

$\partial f / \partial E > 0$ was concluded to be the source of free energy to destabilize this mode, which was identified as a GAE.

Linearized Kinetic-MHD model

$$\gamma\tilde{\psi} + [\tilde{\varphi}, \psi_{eq}] = 0$$

$$\gamma\nabla \cdot (\varrho_{eq}\nabla\tilde{\varphi}) = [\tilde{\psi}, J_{eq}] + [\psi_{eq}, \tilde{J}] - \mathbf{e}_z \cdot \nabla \times (\nabla \cdot \tilde{\mathbf{P}}_h)$$

$$\tilde{\mathbf{P}}_h = \tilde{p}_{\perp h} \mathbf{I} + (\tilde{p}_{\parallel h} - \tilde{p}_{\perp h}) \mathbf{e}_{\parallel} \mathbf{e}_{\parallel}$$

Dispersion relation

$$-\gamma^2 \int d^3x \rho \boldsymbol{\xi} \cdot \boldsymbol{\xi}^* = -\frac{1}{2} \int d^3x \boldsymbol{\xi}^* \cdot \mathbf{F}(\boldsymbol{\xi}) + \frac{1}{2} \int d^3x \boldsymbol{\xi}^* \cdot \nabla \cdot \tilde{\mathbf{P}}_h = \delta W_{MHD} + \delta W_h$$

$\tilde{\mathbf{P}}_h$ is obtained in terms of moments of the perturbed fast ion distribution function.

In the limit $\beta_h \ll \beta_i$, $\delta W_h \ll \delta W_{MHD}$ and only the imaginary part of δW_h matters.

Perturbative approach

To zeroth order, fast particles are neglected $\rightarrow$ rigid-shift solution

To first order, hot particle corrections are considered $\rightarrow \delta W_h$ is evaluated using the zeroth order rigid-shift solution.

The dispersion relation takes on the form

$$\omega^2 = \omega_0^2 - 2i\omega_0\gamma_\eta + \text{Im}(\delta\hat{W}_h)$$

Let $\text{Im}(\delta\hat{W}_h) = i\omega_0^2\lambda_h + \mathcal{O}\left(\frac{\gamma^2}{\omega_0^2}\right)$, $\omega = \omega_0 + \delta\omega$, and $\delta\omega = \mathcal{O}(\gamma/\omega_0)$. Then:

$$\delta\omega = -i\gamma_\eta + i\frac{1}{2}\omega_0\lambda_h = i\gamma$$

Instability requires $\lambda_h > \frac{2\gamma_\eta}{\omega_0}$.

Fast particle δW_h

$$\delta W_h = -\frac{1}{2} \int d^3x d^3v \tilde{L}^* \tilde{f}_h$$

where $\tilde{L} \simeq -(mv_{\parallel}^2 + \mu_{\perp} B) \xi \cdot \kappa$ is the perturbed Lagrangian for fast particle motion. Now:

$$\delta W_h(\omega) = \delta W_1 + \delta W_2(\omega)$$

$$\delta W_2 = -\frac{2\pi^2 c}{Zem^2} \sum_{\sigma} \int dP_{\varphi} d\mathcal{E} d\mu_{\perp} \tau_t \left(\omega \frac{\partial f_h}{\partial \mathcal{E}} - n \frac{\partial f_h}{\partial P_{\varphi}} \right) \sum_{p=-\infty}^{+\infty} \frac{|\Upsilon_p|^2}{\omega + p\omega_t}$$

$$\Upsilon_p(\mathcal{E}, \mu, P_{\varphi}) = \langle \tilde{L} \exp(ip\omega_t \tau) \rangle \quad \text{and} \quad \Upsilon_1 = i \frac{\epsilon^2}{q(r)^2} \mathcal{E} (2 - \Lambda) \frac{\xi}{r} \langle \sin(\vartheta) \sin(\tau\omega_t) \rangle$$

Resonant Interaction

$$\Lambda = \frac{\mu B \phi_0}{\varepsilon}, \quad 0 \leq \Lambda \leq 1 + \epsilon$$

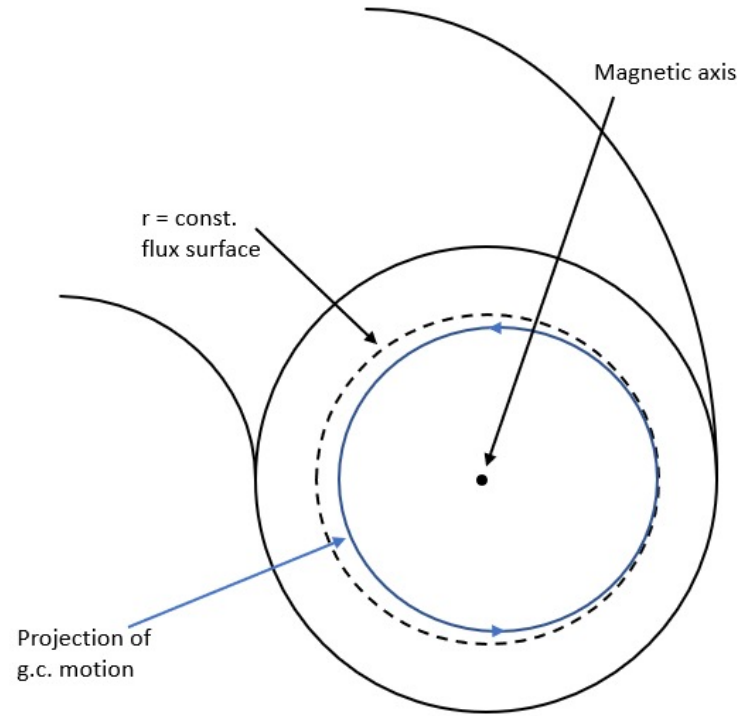
$$\tau_{t/b} = \oint d\tau$$

$$\omega_t = \frac{2\pi}{\tau_t} = \frac{v\kappa}{2R_0q} (2\epsilon)^{1/2} \frac{\pi}{\mathcal{K}(1/\kappa^2)}$$

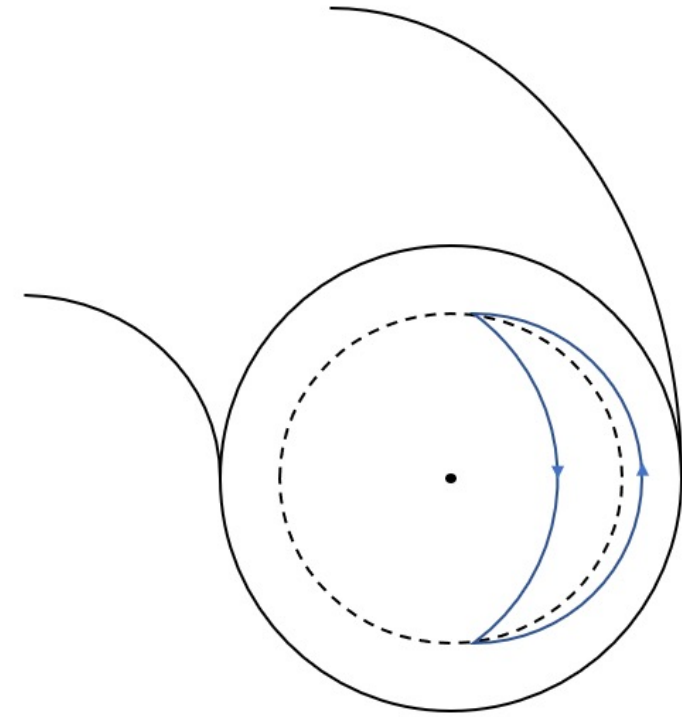
$$\omega_b = \frac{2\pi}{\tau_b} = \frac{v}{2\sqrt{2}R_0q} (\epsilon)^{1/2} \frac{\pi}{\mathcal{K}(\kappa^2)}$$

$$\kappa^2 = \frac{1}{2} + \frac{1}{2\epsilon} (1 - \Lambda)$$

Fast ions trajectories



(a) $\Lambda \leq 1 - \epsilon$



(b) $1 - \epsilon < \Lambda \leq 1 + \epsilon$

$\partial f_h / \partial E > 0$ required for instability

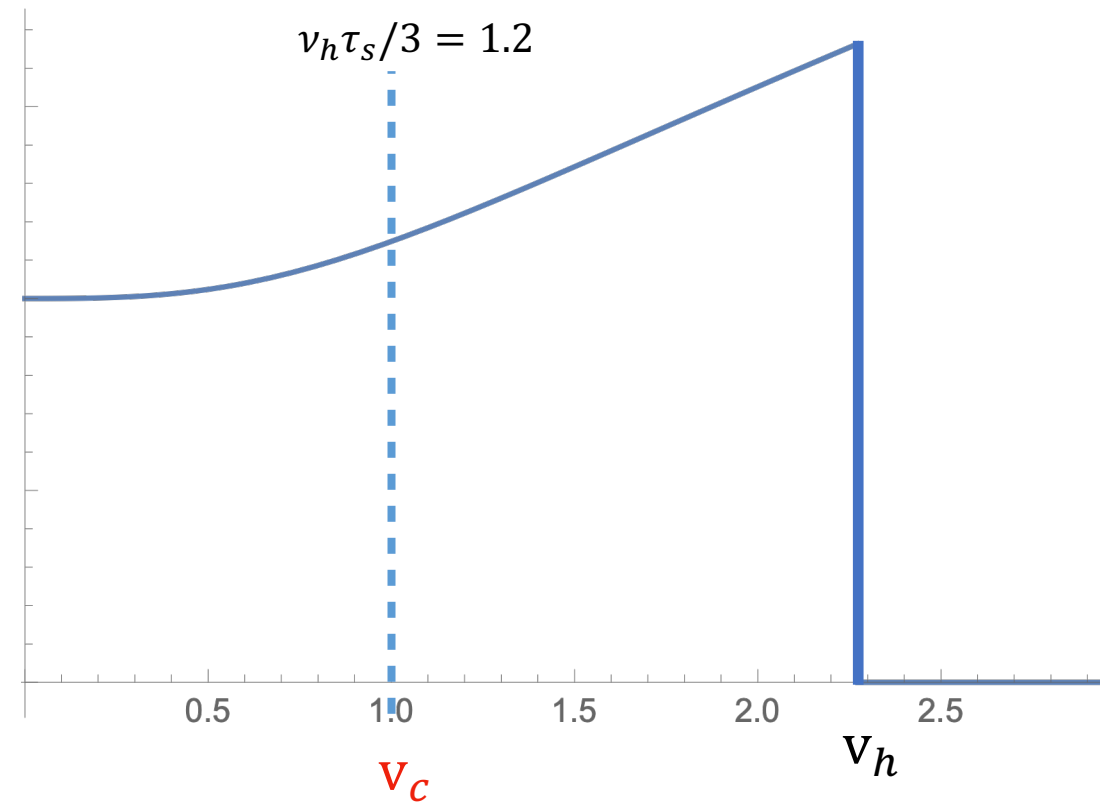
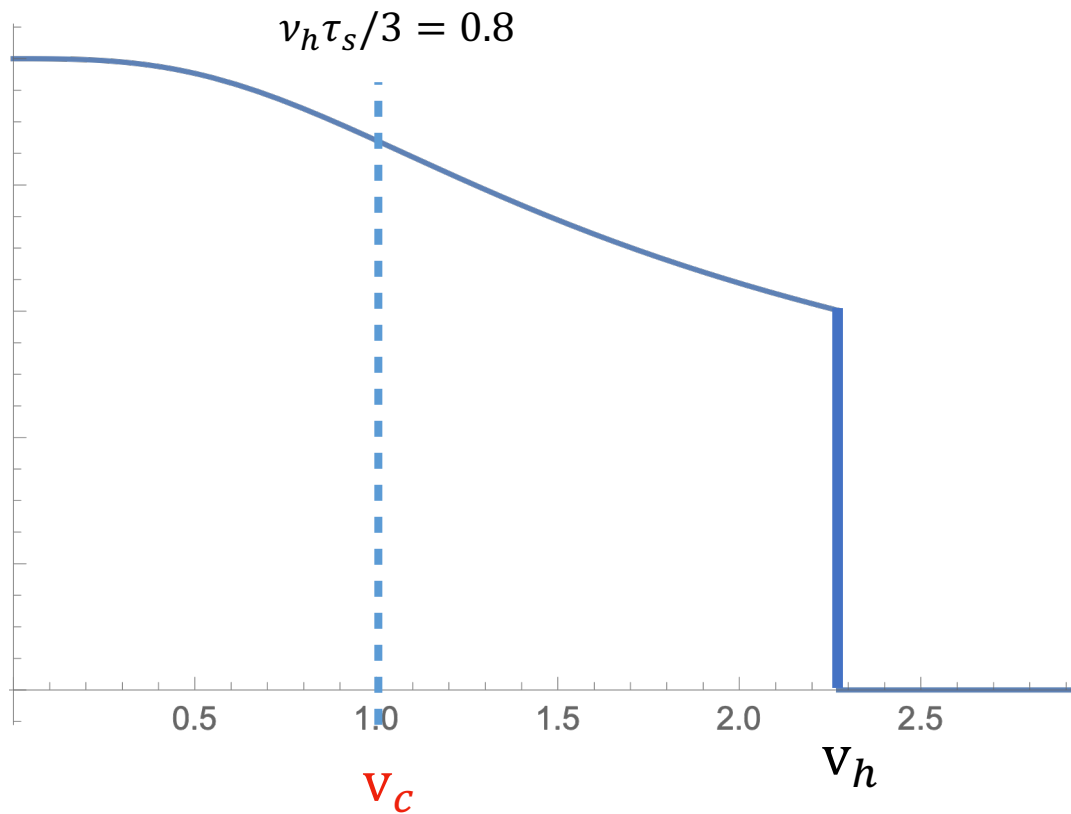
Fokker-Plank equation for fast particle distribution function:

$$\frac{\partial f_h}{\partial t} - \frac{1}{\tau_s v^2} \frac{\partial}{\partial v} [(v^3 + v_c^3) f_h] + \nu_h f_h = S_0 \delta(v - v_h)$$

ν_h : fast particle loss frequency; τ_s : slowing down time

Positive slope obtained for $\frac{\nu_h \tau_s}{3} > 1$.

Solution of F.-P. equation for different values of $\nu_h \tau_s / 3 < 1$



Experimental evidence

M.A. Van Zeeland et al., PoP 2021:

Beam Modulation and Bump-on-tail Effects On Alfvén Eigenmode Stability in DIII-D

Insurgence of bump-on-tail fast ion distribution modulating NBI with periods 7-30 ms (reported: $\tau_s \sim 50\text{ms}$).

V.G. Kiptily et al., NF 2021

Self modulation of alpha particles production due to sawtooth oscillations in JET D-He³ experiments.

Bump-on-tail distribution for energetic particles when the sawtooth period is relatively short, i.e., $\tau_{saw}/\tau_s < 1$.

Critical instability threshold – numerical estimates

$$\lambda_h > \frac{2\gamma_\eta}{\omega_0} \quad \rightarrow \quad \left(\frac{n_h}{n_i}\right)_{crit} = \frac{2\gamma_\eta}{\alpha\omega_0}$$

$$\lambda_h = \alpha(n_h/n_i); \quad \alpha \sim 10^{-1} - 10^{-2}; \quad \gamma_\eta/\omega_0 \sim 10^{-4}$$

$$\left(\frac{n_h}{n_i}\right)_{crit} \sim 10^{-3} - 10^{-2}$$



CONCLUDING REMARKS

Have we forgotten anything?

Check-list for $n=0$ developing theory

- ☒ Ideal-MHD
- ☐ Resistive-MHD
- ☐ Extended-MHD
- ☒ Hybrid Kinetic-MHD
- ☐ SOL Physics, field-line-tying, special boundary conditions
- ☐ other...

Not only $n=0$ vertical displacements

- Axisymmetric current sheet at the magnetic X-point and along the last closed flux surface may affect the stability of peeling modes
- The nonlinear evolution of peeling-ballooning modes is likely to produce a nonlinear $n=0$ component that resonate at the X-point
- ELM mitigation by vertical kicks
-

Thank you for your attention!

